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Enabling precise predictions for left-right symmetry at colliders

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Abstract We investigate the structure of the minimal leftright symmetric model that enables precise predictions in the gauge, scalar and neutrino sector. We revisit the complete set of mass spectra and mixings for the charged and neutral gauge bosons, would-be-Goldstones and gauge fixing, together with the ghost Lagrangian. In the scalar sector, we analytically re-derive all the massive states with mixings and devise a non-trivial physical input scheme, expressing the model couplings in terms of masses and mixing angles. Fermion couplings are also determined in closed form, including the Dirac mixing in the neutrino sector, evaluated explicitly using the Cayley-Hamilton theorem. These analytic developments are implemented in a comprehensive FeynRules model file. We calculate the one loop QCD corrections and provide a complete UFO file for NLO studies, demonstrated on relevant hadron-collider benchmarks. We provide various restricted variants of the model file with different gauges, massless states, neutrino hierarchies and parity violating $g_L \neq g_R$ gauge couplings.

1 Introduction

Understanding the microscopic nature of forces and the spontaneous origin of mass are the two core attributes of the Standard Model (SM) [1]. Neutrino oscillations have now firmly established that neutrinos are massive, in contrast to the SM, and recent significant experimental progress deter-

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mined their mixing properties to unprecedented precision. Still, the nature of neutrino mass (Dirac vs. Majorana) and even more importantly its origin, remain an unsolved mystery of particle physics. Contrary to charged fermions, whose origin is tightly connected to the Higgs mechanism, as the LHC data keeps confirming [2,3], we do not know which theoretical framework is responsible for neutrino mass generation. Another unsolved puzzle in the SM is the glaring parity asymmetry of weak interactions. With the enduring SM experimental success, we got used to the chiral nature of weak interactions and are taking it for given, but the reason behind it remains as unclear as it was at the time of its discovery.

Left-right (LR) symmetry addresses both of these shortcomings within a single framework by postulating parity invariance of weak interactions via interchangeable $SU(2)_L \times SU(2)_R$ local symmetries, featuring new righthanded gauge bosons W_R and Z_R . The need for anomaly cancellation automatically brings in three right-handed neutrinos $N_{i=1,2,3}$, whose mass is tied to the spontaneous breaking scale of $SU(2)_R$. Moreover, the U(1) charge assignment becomes very simple and parity symmetric (or vector-like) and is given by $U(1)_{B-L}$.

Gauging lepton number indicates that the Higgs sector may (and in fact does) lead to lepton number violation (LNV), if neutrinos couple to the Higgs with a Majorana-type Yukawa coupling. The original works addressed the issue of parity restoration [4,5] with soft breaking, but it was soon realized that spontaneous breaking works with both doublets [6] and $\Delta_{L,R}$ triplets [7–10]. The latter option allows for Majorana mass terms for heavy and light neutrinos, which led to the celebrated seesaw mechanism [11,12] and features LNV. We shall refer to it as the minimal left-right symmetric model (LRSM). Unlike GUT-inspired seesaw scenarios [13,14], the LRSM scale need not be much above the elec-

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troweak scale, since the interactions do not mediate proton decay.

Indeed, prior to the advent of LHC, the most stringent constraints on the LRSM scale came from flavor physics. The point here is that, as usual in gauge theories, flavor asymmetry comes from the Yukawa sector when the bi-doublet ϕ couples to quark doublets $Q_{L,R}$. When LR symmetry, in either \mathcal{P} or \mathcal{C} form, is imposed on the Lagrangian, Yukawa couplings are constrained to be nearly hermitian or symmetric. This implies similar flavor mixings in the quark sector $V_R^q \sim V_L^q$ and has led to significant constraints since the Tevatron era [10,15–17] that pushed the scale beyond direct detection capabilities of colliders for some time. The increase of center-of-mass energy at the LHC suffices to probe the scales beyond the flavor constraints [18–26], which spurred a number of collider studies, see reviews in [27–30].

Arguably the most fundamental signal one can look for is the Keung-Senjanović(KS) [31] production of heavy Majorana neutrinos through W_R -mediated gauge interactions. With even a tiny amount of data, the recast of leptoquark searches led to a significant bound [32] that was extensively characterized across the $M_{W_R} - m_N$ plane [33] from well separated $\ell^{\pm}\ell^{\pm}jj$ to boosted neutrino jets [34,35] with a displaced vertex [33, 36] and finally invisible N [33]. (Some of these bounds are relaxed if CKM is different in leftand right-handed interactions [37–39].) With enough data it may be possible to characterize the chirality [40,41], disambiguate production channels [42], look for CP phases in case of degeneracy [43] and include top final states [44]. At even smaller masses, decays of hadrons and mesons become important when a significant amount of decaying states is available [45–47]. At very high W_R masses, off-shell production takes place [48]. Looking into the future, the reach of an FCC-hh was estimated to be around 30 TeV [49], in agreement with other studies [48, 50-52]. Similarly, the potential sensitivity of displaced vertices at lepton collider can probe the M_{W_R} up to ~ 26 TeV at FCC-ee and ~ 70 TeV for a muon collider [53].

At present, many experimental searches by CMS and ATLAS are targeting W' and Z' resonances. Here we focus on W_R , which is lighter than Z_R and thus sets the most stringent constraint on the LR breaking scale and we only review the most recent experimental updates. The KS channel [54,55] searches are looking for resolved as well as boosted (but prompt) topologies of N decays. In the low m_N regime, the $W_R \rightarrow \ell N$ channel is equivalent to $W' \rightarrow \ell v$, because N becomes long-lived enough to decay outside of the detector. One can then recast both leptonic (e, μ) [56,57] as well as τ final states [58–60]. In the LR symmetric case, the CKM in the left and right-handed sectors are similar and one has direct limits from dijet [61,62] and tb resonance [63,64] searches. The final state topologies here are clearly independent of m_N and so is the resulting limit on M_{W_R} , apart from the slight

suppression of the Br when N channels open up. Finally, W_R can decay to SM gauge bosons W, Z and Higgs h, leading to further dedicated searches [65,66].

The imposition of LR-symmetry has important ramifications in the leptonic sector, such as $\mu \rightarrow e\gamma$, which was the original motivation for seesaw [11]. When the triplet Yukawa couplings are restricted to a near-hermitian or symmetric form. Dirac and Majorana couplings become directly related. In the case of C [67] this connection allows one to calculate the Dirac mass in terms of Majorana masses, via a square root of a matrix. The new analytic solution found in this work gives a direct insight into the Dirac-Majorana connection from colliders, where Majorana masses may be measured, and connect it to any Dirac-mediated process. These processes appear either at colliders [67–70], nuclear $0\nu 2\beta$ [12,71–74], electron EDMs or radiative decays, and may be relevant for dark matter searches [75-77]. The connection becomes rather involved in the case of \mathcal{P} [78,79] and one can resort to numerical algorithms [80] to invert it.

The primary focus of LHC studies has been on the gauge sector, looking for a massive TeV scale W_R resonance in the s-channel and the associated Majorana neutrino [20, 28, 32, 33, 42]. On the other hand, the Higgs sector is of fundamental importance for experimentally establishing the spontaneous mass origin of heavy neutrinos from Δ_R Yukawa coupling to N. The scalar sector was introduced in [6,7] and analyzed in follow-up studies [81-85] to the more recent works [86-89] with emphasis on perturbativity [90–92], vacuum stability [93–95] and gravitational waves [96,97]. Our approach is to start from physical quantities, masses and mixing angles and use them to compute the potential couplings. We therefore revisit and re-derive the minimization conditions and compute the spectrum to give a closed form solution for the potential parameters in terms of physical inputs.

There are two simplifying assumptions that we make in the entire analysis. The first regards the vev of Δ_L , which has to be small for phenomenological reasons, limited to $v_L \leq$ GeV by the ρ parameter. For the most part of our discussion we set $v_L = 0$, which simplifies the diagonalization of the scalar mass spectra without a significant impact on the eigenvalues and phenomenology. The other assumption regards the two quartic couplings of the bi-doublet, which we take to be related by $\lambda_2 = 2\lambda_3$. This removes a part of the mixing between the scalar and pseudo-scalar components and drastically simplifies the neutral mass matrix, thus allowing for analytic diagonalization without significant phenomenological impact of additional CP violation in the heavy scalar sector.

The most fundamental conceptual goal in the Higgs sector, at least in relation to neutrino mass origin, is to prove the spontaneous origin of the heavy N masses. In this mechanism the decay rate of $\Delta_R \rightarrow NN$ is proportional to m_N , which can be measured kinematically, just like $\Gamma_{h\to f\bar{f}} \propto m_f^2$ in the SM, where m_f are already known. The two scalars h and Δ_R can mix and we get both $h \to NN$ from the 125 GeV resonance [98] and/or a sizeable production of $\Delta_R \to NN$ at the unknown m_{Δ} [93]. Such signals of Majorana Higgses would finalize the program of determining the neutrino mass origin and complete the microscopic picture of massive neutrinos.

The rest of the Higgs sector is also conceptually interesting but phenomenologically challenging [89,90]. The scalar H, pseudoscalar A and singly charged H^+ from the bi-doublet ϕ can all mediate FCNCs at tree level and therefore have to be quite above the TeV scale. They are unlikely to be accessible at the LHC but might be visible at future colliders [92]. If W_R were light enough to be seen at the LHC, the entire Δ_L would be somewhat split [90] and would appear mainly in cascade decays [99]. The doubly charged Δ_R^{++} may be produced in association with W_R [100] and would lead to additional LNV final states [101, 102] or long-lived dE/dx signatures [103].

Clearly, there is plenty of rich and conceptually relevant phenomenology to be studied in the context of the LRSM, at colliders, low energies (T2K [47], SHiP [104] and FASER [105]) and in cosmology. To ease such studies it is desirable to have a complete analytical model solution and possibly a consistent implementation in a model file. This is precisely the purpose of this work, where we solve analytically the model and provide an implementation within Feyn-Rules (FR) [106, 107] and related UniversalFeynmanOutput (UFO) versions [108, 109].

We collect and re-derive all the parts of the LRSM, starting from covariant derivatives and gauge eigen-states, together with would-be-Goldstones and ghosts, implementing a switch between the Feynman and unitary gauge. A study of renormalization was performed in the gauge sector [110] and in the Higgs sector [88] at one loop. An earlier tree-level implementation of the LRSM can be found in [111] with arbitrary Yukawa couplings and no mixing in the scalar sector. In the context of seesaw, there are available model files at NLO for type I seesaw/singlets [112], type II [113] and type III [114], as well as effective W'/Z' model [115,116] and the Weinberg operator [117].

In this work we pay special attention to the scalar sector and derive anew analytic expressions for masses and mixings, providing a new physical input scheme and perturbativity constraints. We take into account the symmetry-imposed flavor structure in the quark sector, leading to symmetric CKM mixings and calculable Yukawa matrices of the extended Higgs sector to fermions. In the neutrino sector we focus in detail on the Majorana–Dirac connection and implement the heavy-light neutrino mixing by devising an explicit closed form solution for the root(s) using the Cayley–Hamilton theorem. This is the first time that a complete and calculable heavy-light neutrino mixing is implemented in a model file, including the current neutrino oscillation data. Finally, we obtain the QCD one loop counter-terms and provide, for the first time, the complete LRSM at NLO, which enhances and simplify future collider studies. We demonstrate the use of the model file on a number of benchmarks that exemplify these developments and quantify, in some cases for the first time, the reduction of uncertainties in LNV signals at the LHC and beyond.

The paper is organized as follows: in Sect. 2 we introduce the minimal LRSM with the relevant fields. We discuss each of its sectors in a separate sub-section, going from gauge bosons and would-be-Goldstones in Sect. 2.1 to scalars Sect.2.2 (including the inversion of couplings Sect.2.3), fermions Sect.2.4 and finishing with ghosts in Sect.2.5. The Sect. 3 is dedicated to the variant of the model with different gauge couplings. The core of the implementation of the model file is laid out in Sect. 4 and the QCD NLO corrections are computed in Sect. 4.2. Some phenomenologically relevant benchmarks are given in Sect. 5, while the summary and outlook is done in Sect. 6. The Appendix A explains the Cayley–Hamilton approach used to obtain an analytic closed form of the square root of a matrix.

2 The minimal left-right symmetric model

The minimal left-right symmetric model is based on a parity symmetric gauge group

$$G_{LR} = SU(3)_c \otimes SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}, \qquad (1)$$

where some discrete symmetry interconnects the two $SU(2)_{L,R}$ gauge groups. These two groups are assumed to have the same gauge couplings at some energy scale. We shall be somewhat agnostic about this and first take $g_L = g_R = g$ for simplicity. Later on we generalize to $g_L \neq g_R$, which turns out not to be a very drastic change. The gauge coupling of the remaining $U(1)_{B-L}$ is g'.

In the minimal LRSM, parity and the gauge symmetry are broken spontaneously in the scalar sector, featuring a bidoublet $\phi = (1, 2, 2, 0)$ and two triplets $\Delta_L = (1, 3, 1, 2)$, $\Delta_R = (1, 1, 3, 2)$ under G_{LR} ,

$$\phi = \begin{pmatrix} \phi_1^{0*} \ \phi_2^+ \\ \phi_1^- \ \phi_2^0 \end{pmatrix}, \qquad \Delta_{L,R} = \begin{pmatrix} \frac{\Delta^+}{\sqrt{2}} \ \Delta^{++} \\ \Delta^0 \ -\frac{\Delta^+}{\sqrt{2}} \end{pmatrix}_{L,R}.$$
 (2)

The spontaneous symmetry breaking, which leaves unbroken only the electric charge generator

$$Q = T_{3L} + T_{3R} + \frac{B - L}{2},\tag{3}$$

follows from the vevs of the neutral components,

$$\langle \phi \rangle = \begin{pmatrix} v_1 & 0\\ 0 & -e^{i\alpha}v_2 \end{pmatrix}, \qquad \langle \Delta_{L,R} \rangle = \begin{pmatrix} 0 & 0\\ v_{L,R} & 0 \end{pmatrix}, \qquad (4)$$

where

$$v^2 = v_1^2 + v_2^2$$
 and $0 \le \tan \beta = \frac{v_2}{v_1} < 1$, (5)

and we assumed $v_1 > v_2$ without loss of generality.

These vevs are quite hierarchical

$$v_L \lesssim \text{GeV}, \qquad v = 174 \text{ GeV}, \qquad v_R \gtrsim \text{TeV}, \quad (6)$$

and generate the masses of all the states: the gauge and Higgs bosons as well as fermions. The SM-like gauge bosons W_L , Z are on the order of v and so is the Higgs boson h. The RH gauge bosons W_R , Z_{LR} are heavy, on the order of v_R and so are most of the remaining scalars in $\Delta_{L,R}$ and the mass eigenstates H, A, H^{\pm} from the bi-doublet. Of course, the photon remains massless, while the Goldstones are initially massless and then get their unphysical masses from gauge fixing.

The fermions come in LR symmetric doublet representations

$$Q_{L,R} = \begin{pmatrix} u \\ d \end{pmatrix}_{L,R}, \qquad \qquad L_{L,R} = \begin{pmatrix} v \\ \ell \end{pmatrix}_{L,R}, \tag{7}$$

with the usual three family copies. The appealing feature of this setup is that the three RH neutrinos v_R are required by anomaly cancellation. In the LRSM, all the fermion masses come from spontaneous breaking of the two SU(2) groups, meaning their masses are a product of Yukawa couplings and vevs. For O(1) Yukawa couplings, the Dirac masses are on the order of v, the Majorana mass for v_R is $O(v_R)$, while the light neutrinos v_L get their Majorana mass as a combination of v^2/v_R and v_L terms (combination of type I and type II seesaw).

In the following Sect. 2.1 we compute the gauge boson masses and mixings and derive the required gauge fixing terms. Then we focus on the scalar sector, which includes the Goldstones and the physical states. We shall calculate the mass spectra, define an input scheme and compute the parameters of the potential in terms of masses and mixings.

2.1 Gauge sector

Gauge boson masses and mixings. Let us begin with the calculation of the gauge boson mass spectrum in the LRSM. The canonical kinetic terms for complex scalar fields are

$$\mathcal{L}_{\rm kin} = |D\phi|^2 + |D\Delta_L|^2 + |D\Delta_R|^2 , \qquad (8)$$

with

$$D\phi = \partial\phi - ig\left(A_L\phi - \phi A_R\right), \qquad (9)$$

$$D\Delta_i = \partial \Delta_i - ig \left[A_i, \Delta_i\right] - ig' B\Delta_i , \qquad (10)$$

where i = L, *R* and the gauge and Lorentz indices are suppressed. The gauge bosons fields in the flavor basis are $A_i = A_i^a \sigma^a/2$, $A_i^{\pm} = (A_i^1 \mp i A_i^2)/\sqrt{2}$, with implied sum over the repeated indices a = 1, 2, 3.

The signs of the gauge coupling terms in the covariant derivatives (9) and (10) are not uniquely defined, as is the case for the SM [118]. The convention above is chosen to match the SM-like decoupling limit, such that when $\tan \beta \rightarrow 0$, the ϕ_1 takes on the role of the SM Higgs and has the usual covariant derivative. One also has to keep a consistent choice of signs between the bi-doublet ϕ and Δ_R . None of these considerations affect the gauge boson mass calculation, they only come into play when we consider the gauge fixing that has to remove the mixed generic $V\partial\phi$ terms.

Gauge boson masses and mixings. Covariant derivatives in (9) and (10) generate the gauge boson mass matrices for charged and neutral gauge bosons

$$\begin{pmatrix} A_{L}^{-} A_{R}^{-} \end{pmatrix} M_{W_{LR}}^{2} \begin{pmatrix} A_{L}^{+} \\ A_{R}^{+} \end{pmatrix}$$

$$= \begin{pmatrix} W_{L}^{-} W_{R}^{-} \end{pmatrix} \begin{pmatrix} M_{W_{L}}^{2} & 0 \\ 0 & M_{W_{R}}^{2} \end{pmatrix} \begin{pmatrix} W_{L}^{+} \\ W_{R}^{+} \end{pmatrix}.$$
(11)

The hermitian mass matrix $M_{W_{IR}}^2$ in the gauge basis is

$$M_{W_{LR}}^2 = \frac{g^2}{2} v_R^2 \begin{pmatrix} \epsilon^2 & e^{-i\alpha} \epsilon^2 s_{2\beta} \\ e^{i\alpha} \epsilon^2 s_{2\beta} & 2 + \epsilon^2 \end{pmatrix}, \tag{12}$$

with the short-hands $s_x = \sin(x)$, $c_x = \cos(x)$ and $t_x = \tan(x)$. We introduced the dimensionless ratio of vevs

$$\epsilon = \frac{v}{v_R},\tag{13}$$

which needs to be small on phenomenological grounds.

We compute the eigenvalues and expand them in ϵ

$$M_{W_L} \simeq \frac{gv}{\sqrt{2}}, \qquad \qquad M_{W_R} \simeq gv_R\left(1 + \frac{\epsilon^2}{4}\right).$$
 (14)

We drop higher order corrections of the order of ϵ^4 in both the $M_{W_L}^2$ and $M_{W_R}^2$. Specifically, the ϵ^2 term in $M_{W_R}^2$ is of the same order as $M_{W_r}^2$.

To get from the gauge/flavor to the mass basis

$$\begin{pmatrix} A_L^+\\ A_R^+ \end{pmatrix} = U_W \begin{pmatrix} W_L^+\\ W_R^+ \end{pmatrix},\tag{15}$$

we have the following 2D unitary rotation matrix

$$U_W = \begin{pmatrix} c_{\xi} & s_{\xi} e^{-i\alpha} \\ -s_{\xi} e^{i\alpha} & c_{\xi} \end{pmatrix}, \tag{16}$$

with the mixing angle given by

$$s_{\xi} \simeq \frac{\epsilon^2}{2} s_{2\beta} \simeq \frac{M_{W_L}^2}{M_{W_R}^2} s_{2\beta} \,. \tag{17}$$

The W^- states get rotated with U_W^* , such that

$$U_{W}^{\dagger} M_{W_{LR}}^{2} U_{W} = \begin{pmatrix} M_{W_{L}}^{2} & 0\\ 0 & M_{W_{R}}^{2} \end{pmatrix}.$$
 (18)

We move on to neutral gauge bosons between the following bases

$$\begin{pmatrix} A_{3L} \ A_{3R} \ B \end{pmatrix} M_{Z_{LR}}^{2} \begin{pmatrix} A_{3L} \\ A_{3R} \\ B \end{pmatrix}$$

$$= \begin{pmatrix} A \ Z_{L} \ Z_{R} \end{pmatrix} \begin{pmatrix} 0 & 0 & 0 \\ 0 \ M_{Z}^{2} & 0 \\ 0 & 0 \ M_{Z_{LR}}^{2} \end{pmatrix} \begin{pmatrix} A \\ Z_{L} \\ Z_{R} \end{pmatrix},$$
(19)

where the symmetric mass matrix in the gauge basis is given by

$$M_{Z_{LR}}^2 = \frac{g^2}{2} v_R^2 \begin{pmatrix} \epsilon^2 & -\epsilon^2 & 0\\ -\epsilon^2 & 4 + \epsilon^2 & -4r\\ 0 & -4r & 4r^2 \end{pmatrix}, \qquad r = \frac{g'}{g}.$$
 (20)

This is a real matrix (no CP phases) that does not depend on t_{β} , and its eigenvalues can be expanded in ϵ

$$M_A = 0, \quad M_Z \simeq \frac{gv}{\sqrt{1 + \frac{1}{1 + 2r^2}}},$$
 (21)

$$M_{Z_{LR}} \simeq g v_R \sqrt{2(1+r^2)} \left(1 + \frac{\epsilon^2}{8(1+r^2)^2}\right).$$
 (22)

One way to define the weak mixing angle is through the ratio of gauge boson masses at leading order in ϵ

$$c_w = \frac{M_{W_L}}{M_{Z_L}}\Big|_{\mathcal{O}(\epsilon^0)},\tag{23}$$

which in turn determines the ratio

$$r = \frac{s_w}{\sqrt{c_{2w}}} \simeq \sqrt{\frac{M_{W_L}^2}{2M_{W_L}^2 - M_{Z_L}^2} - 1} \simeq 0.63.$$
(24)

From (22) one then recovers the prediction

$$\frac{M_{Z_{LR}}}{M_{W_R}} \simeq \sqrt{\frac{2c_w^2}{c_{2w}}} \left(1 + \frac{\epsilon^2}{8} \frac{c_{2w}^2}{c_w^4} \right) \simeq 1.67.$$
(25)

With these definitions, the real orthogonal matrix O_Z that takes us from the flavor to the mass basis,

$$\begin{pmatrix} A_{3L} \\ A_{3R} \\ B \end{pmatrix} = O_Z \begin{pmatrix} A \\ Z_L \\ Z_R \end{pmatrix},$$
(26)

is found at second order in ϵ to be

$$O_{Z} = \begin{pmatrix} s_{w} & -c_{w} & 0\\ s_{w} & s_{w}t_{w} & -\frac{\sqrt{c_{2w}}}{c_{w}}\\ \sqrt{c_{2w}} & \sqrt{c_{2w}}t_{w} & t_{w} \end{pmatrix} + \frac{\epsilon^{2}}{4} \begin{pmatrix} 0 & 0 & \frac{c_{2w}^{3/2}}{c_{w}^{3}} \\ 0 & -\frac{c_{2w}^{2}}{c_{w}^{3}} - \frac{c_{2w}^{3/2}t_{w}}{c_{w}^{3}} \\ 0 & \frac{c_{2w}^{3/2}t_{w}}{c_{w}^{4}} - \frac{c_{2w}^{2}t_{w}}{c_{w}^{4}} \end{pmatrix}.$$
(27)

This concludes the analysis of the gauge boson spectrum.

Goldstones and gauge fixing. To complete the description of massive gauge bosons with spontaneous symmetry breaking, we need to address the would-be-Goldstone (wbG) boson sector. Prior to gauge fixing, these appear as massless states of the scalar mass matrices. In the SM there is only one neutral and one charged wbG, however in theories with an extended scalar sector, like the LRSM, there are more such states and they typically mix. Thus we get degenerate zeroes and additional freedom of rotation between massless states.

In the case of the LRSM, we are dealing with two charged wbG modes for the $W_{L,R}$ and two neutral ones for Z, Z_{LR} gauge bosons, thus this freedom comes as one mixing angle for each sector. These angles are not arbitrary, because the proper wbG mode has to couple derivatively and in proportion to the corresponding gauge boson mass, i.e. $M_V(V\partial\phi)$. In other words, one needs to isolate the proper wbG mode for each gauge boson. Eventually, we will add the gauge fixing terms in the mass basis, which cancel the derivative mixed terms and generate the usual gauge dependent propagators for vectors and wbGs.

We discuss first the singly charged states and derive their mass matrix. It comes from the mixed second derivatives of the potential $d^2 V/(d\phi_i^+ d\phi_j^-)$, evaluated at its minimum. The minimization is described in the scalar sector section below. We use the following basis

$$\begin{pmatrix} \phi_1^- \ \phi_2^- \ \Delta_R^- \end{pmatrix} M_+^2 \begin{pmatrix} \phi_1^+ \\ \phi_2^+ \\ \Delta_R^+ \end{pmatrix}$$

$$= \begin{pmatrix} \varphi_L^- \ H^- \ \varphi_R^- \end{pmatrix} \begin{pmatrix} 0 & 0 & 0 \\ 0 & m_{H^+}^2 & 0 \\ 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} \varphi_L^+ \\ H^+ \\ \varphi_R^+ \end{pmatrix},$$

$$(28)$$

where we dropped the Δ_L^{\pm} , because of the negligible v_L .

The singly charged mass matrix and its only nonzero eigenvalue are given by

$$M_{+}^{2} = \alpha_{3} v_{R}^{2} \begin{pmatrix} \frac{s_{\beta}^{2}}{c_{2\beta}} & -e^{-i\alpha \frac{t_{2\beta}}{2}} & -\epsilon e^{-i\alpha \frac{s_{\beta}}{\sqrt{2}}} \\ -e^{i\alpha \frac{t_{2\beta}}{2}} & \frac{c_{\beta}^{2}}{c_{2\beta}} & \epsilon \frac{c_{\beta}}{\sqrt{2}} \\ -\epsilon e^{i\alpha \frac{s_{\beta}}{\sqrt{2}}} & \epsilon \frac{c_{\beta}}{\sqrt{2}} & \epsilon^{2} \frac{c_{2\beta}}{2} \end{pmatrix}, \quad (29)$$

$$m_{H^+} \simeq \sqrt{\frac{\alpha_3}{c_{2\beta}}} v_R \left(1 + \epsilon^2 \frac{c_{2\beta}^{3/2}}{4} \right) \,. \tag{30}$$

As expected, there are two zero modes in (28) that are eaten up by the SM-like W_L and the new W_R .

One can diagonalize M_+^2 in (29) by two subsequent rotations. However, this is not enough: after this diagonalization and after performing the U_W rotation among $W_{L,R}$, the mixed derivative terms, coming from (9) and (10), are still off-diagonal. In particular, while both $W_{L,R}^- \partial H^+$ terms vanish as they should, the other four combinations $W_{L,R}^- \partial \varphi_{L,R}^+$ between incompletely rotated $\varphi_{L,R}^+$ are all there. We need to perform one final rotation between wbGs to get rid off the two off-diagonal terms, which finally extracts the canonical would-be-Goldstone modes.

Combining all the rotations, we get the complete transformation of the charged scalars

$$\begin{pmatrix} \phi_1^+ \\ \phi_2^+ \\ \Delta_R^+ \end{pmatrix} = U_+ \begin{pmatrix} \varphi_L^+ \\ H^+ \\ \varphi_R^+ \end{pmatrix}, \tag{31}$$

where the unitary matrix U_+ is given by three subsequent rotations and can also be expanded in ϵ ,

$$U_{+} = \begin{pmatrix} c_{\beta} & -e^{-i\alpha}s_{\beta} & 0\\ e^{i\alpha}s_{\beta} & c_{\beta} & 0\\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0\\ 0 & -1 + \frac{1}{4}\epsilon^{2}c_{2\beta}^{2} & \frac{\epsilon c_{2\beta}}{\sqrt{2}}\\ 0 & -\frac{\epsilon c_{2\beta}}{\sqrt{2}} & -1 + \frac{1}{4}\epsilon^{2}c_{2\beta}^{2} \end{pmatrix}$$

$$\times \begin{pmatrix} 1 - \frac{1}{4}\epsilon^{2}s_{2\beta}^{2} & 0 & \frac{\epsilon s_{2\beta}}{\sqrt{2}}\\ 0 & 1 & 0\\ -\frac{\epsilon s_{2\beta}}{\sqrt{2}} & 0 & 1 - \frac{1}{4}\epsilon^{2}s_{2\beta}^{2} \end{pmatrix} \simeq \begin{pmatrix} c_{\beta} & e^{-i\alpha}s_{\beta} & 0\\ e^{i\alpha}s_{\beta} & -c_{\beta} & 0\\ 0 & 0 & -1 \end{pmatrix}$$

$$+ \frac{\epsilon}{\sqrt{2}} \begin{pmatrix} 0 & 0 & e^{-i\alpha}s_{\beta}\\ 0 & 0 & c_{\beta}\\ e^{i\alpha}s_{2\beta} & -c_{2\beta} & 0\\ -4e^{i\alpha}s_{\beta}c_{\beta}^{4} & c_{\beta}c_{2\beta}^{2} & 0\\ 0 & 0 & 1 \end{pmatrix}, \qquad (32)$$

which is unitary up to $\mathcal{O}(\epsilon^2)$.

After performing all the rotations with U_W and U_+ , we end up with canonical wbGs that have derivative couplings proportional to the corresponding gauge boson masses

$$\mathcal{L}_{\rm kin} \ni -i \sum_{i=L,R} M_{W_i} W_i^- \partial \varphi_i^+ \,. \tag{33}$$

The final step is to add the gauge fixing terms, which come in the usual form, written in the mass basis

$$\mathcal{L}_{\rm gf} \ni -\sum_{i=L,R} \frac{1}{\xi_{W_i}} F_i^+ F_i^-,$$
 (34)

$$F_i^+ = \partial W_i^+ + i\xi_{W_i} M_{W_i} \varphi_i^+ \,. \tag{35}$$

After expanding all the terms in \mathcal{L}_{gf} and integrating by parts, the mixed $V \partial \varphi$ terms in (33) cancel away. The remaining terms in (34) give the standard R_{ξ} propagators for the *W*'s and wbGs.

In the neutral sector, the procedure of setting up the wbG rotations and fixing the gauge goes along the same lines. The major additional complication is that there are more neutral states that mix and computing their eigenvalues and mixings becomes a bit more tedious. However, it is fairly simple to isolate the wbG modes away from the massive scalars and complete all the rotations in the gauge-wbG sector. The following field substitution in (2) gets the job done

$$\phi_{10} = v_1 + \frac{1}{\sqrt{2}} \left(c_\beta \varphi_{10} + s_\beta e^{-i\alpha} \varphi_{20} \right) \,, \tag{36}$$

$$\phi_{20} = -v_2 e^{i\alpha} + \frac{1}{\sqrt{2}} \left(c_\beta \varphi_{20} - s_\beta e^{i\alpha} \varphi_{10} \right). \tag{37}$$

With this rotation, the $\Im(\varphi_{10})$ and $\Im(\Delta_R^0)$ decouple from the other neutral scalars and they both have a zero eigenvalue. However, their derivative couplings are still off-diagonal, so we need to perform one final 2*D* rotation to get to the diagonal neutral wbG states

$$\begin{pmatrix} \Im(\varphi_{10}) \\ \Im(\Delta_R^0) \end{pmatrix} = O_{\varphi} \begin{pmatrix} \varphi_L^0 \\ \varphi_R^0 \end{pmatrix},$$
 (38)

$$O_{\varphi} = I_2 - \frac{i\sigma_2}{2(1+r^2)}\epsilon - \frac{I_2}{8(1+r^2)^2}\epsilon^2.$$
 (39)

Here, I_2 is the 2D identity matrix and σ_2 is the usual Pauli matrix. The resulting mixed derivative terms are now diagonal and proportional to the neutral gauge boson masses

$$\mathcal{L}_{\rm kin} \ni -\sum_{i=L,R} M_{Z_i} Z_i \,\partial \varphi_i^0. \tag{40}$$

Again, they get cancelled by the terms that complete the gauge fixing

$$\mathcal{L}_{\rm gf} \ni -\sum_{i=L,R} \frac{1}{2\xi_{Z_i}} F_{Z_i}^2 , \quad F_{Z_i} = \partial Z_i + \xi_{Z_i} M_{Z_i} \varphi_i^0, \quad (41)$$

which also give the canonical R_{ξ} propagators for the neutral gauge bosons and wbGs.

The scalar potential of the LRSM comprises all possible bilinear and quartic terms of the scalar fields, obeying the gauge symmetry, and one discrete LR symmetry, either \mathcal{P} or \mathcal{C} :

$$\begin{split} \mathcal{V} &= -\mu_1^2 \left[\phi^{\dagger} \phi \right] - \mu_2^2 \left(\left[\tilde{\phi} \phi^{\dagger} \right] + \left[\tilde{\phi}^{\dagger} \phi \right] \right) \\ &- \mu_3^2 \left(\left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] + \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right] \right) \\ &+ \lambda_1 \left[\phi^{\dagger} \phi \right]^2 + \lambda_2 \left(\left[\tilde{\phi} \phi^{\dagger} \right]^2 + \left[\tilde{\phi}^{\dagger} \phi \right]^2 \right) \\ &+ \lambda_3 \left[\tilde{\phi} \phi^{\dagger} \right] \left[\tilde{\phi}^{\dagger} \phi \right] + \lambda_4 \left[\phi^{\dagger} \phi \right] \left(\left[\tilde{\phi} \phi^{\dagger} \right] + \left[\tilde{\phi}^{\dagger} \phi \right] \right) \\ &+ \rho_1 \left(\left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right]^2 + \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right]^2 \right) \\ &+ \rho_2 \left(\left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] \left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] + \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right] \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right] \right) \\ &+ \rho_3 \left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right] \\ &+ \rho_4 \left(\left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right] + \left[\Delta_R^{\dagger} \Delta_R^{\dagger} \right] \right) \\ &+ \alpha_1 \left[\phi^{\dagger} \phi \right] \left(\left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] + \left[\Phi^{\dagger} \phi \Delta_R^{\dagger} \right] \right) \\ &+ \left(\alpha_2 \left(\left[\tilde{\phi} \phi^{\dagger} \right] \left[\Delta_L^{\dagger} \Delta_L^{\dagger} \right] + \left[\phi^{\dagger} \phi \Delta_R^{\dagger} \right] \right) \\ &+ \beta_1 \left(\left[\phi \Delta_R^{\dagger} \phi^{\dagger} \Delta_L^{\dagger} \right] + \left[\phi^{\dagger} \Delta_L^{\dagger} \phi \Delta_R^{\dagger} \right] \right) \\ &+ \beta_3 \left(\left[\phi \Delta_R^{\dagger} \phi^{\dagger} \Delta_L^{\dagger} \right] + \left[\phi^{\dagger} \Delta_L^{\dagger} \phi \Delta_R^{\dagger} \right] \right), \end{split}$$
(42)

where $\tilde{\phi} = \sigma_2 \phi^* \sigma_2$ and the square brackets imply the trace over field components. In the following, we set for simplicity $\beta_i \simeq 0$, which sets $v_L = 0$ and remains small; this is a technically natural assumption.

The \mathcal{P} or \mathcal{C} discrete symmetries further constrain the coupling constants.

The case of \mathcal{P} . With this choice, the scalars transform as

$$\mathcal{P}: \phi \to \phi^{\dagger}, \qquad \Delta_L \leftrightarrow \Delta_R, \qquad (43)$$

and all the couplings in the potential must be real except for α_2 , which carries a phase $e^{i\delta_2}$.

The minimization conditions can be interpreted as determining the μ_i in terms of the vevs and quartic couplings:

$$\mu_{1}^{2} = 2 \left(\lambda_{1} + s_{2\beta}c_{\alpha}\lambda_{4}\right)v^{2} + \left(\alpha_{1} - \alpha_{3}\frac{s_{\beta}^{2}}{c_{2\beta}}\right)v_{R}^{2}, \qquad (44)$$
$$\mu_{2}^{2} = \left(s_{2\beta}\left(2c_{2\alpha}\lambda_{2} + \lambda_{3}\right) + \lambda_{4}\right)v^{2} + \frac{1}{2c_{\alpha}}\left(2c_{\alpha+\delta_{2}}\alpha_{2} + \alpha_{3}\frac{t_{2\beta}}{2c_{\alpha}}\right)v_{R}^{2}, \qquad (45)$$

$$\mu_3^2 = \left(\alpha_1 + \left(2c_{\alpha+\delta_2}\alpha_2 s_{2\beta} + \alpha_3 s_{\beta}^2\right)\right)v^2 + 2\rho_1 v_R^2.$$
(46)

where v_L is neglected, since $v_L \ll v \ll v_R$. In fact, one has

$$v_L = \frac{\epsilon^2 v_R}{\left(1 + t_\beta^2\right) (2\rho_1 - \rho_3)} \left(-\beta_1 t_\beta \cos(\alpha - \theta_L) + \beta_2 \cos(\theta_L) + \beta_3 t_\beta^2 \cos(2\alpha - \theta_L)\right).$$

$$(47)$$

The derivative with respect to α connects it to the only CP phase δ_2 :

$$\alpha_2 s_{\delta_2} = \frac{s_\alpha}{4} \left(\alpha_3 t_{2\beta} + 4 \left(\lambda_3 - 2\lambda_2 \right) s_{2\beta} \epsilon^2 \right). \tag{48}$$

The case of C. This acts as

$$\mathcal{C}: \phi \to \phi^T, \qquad \Delta_L \leftrightarrow \Delta_R^*.$$
 (49)

This choice allows for additional complex phases in the potential. In particular, the couplings μ_2 , λ_2 , λ_4 , ρ_4 and β_i are now complex, introducing seven new CP violating phases. The minimization conditions described above also change accordingly.

However, for collider studies, these additional phases do not lead to relevant new features, while spoiling the exact solvability of the scalar spectrum that we achieve below. Therefore, in the present work we retain the single CP phase δ_2 as in the case of \mathcal{P} . Similarly, we can consider the limit of vanishing v_L , and accordingly we set $\beta_i \rightarrow 0$.

2.2.1 Doubly charged Δ_R^{++} and the entire Δ_L triplet

In the $v_L \rightarrow 0$ limit, the doubly charged Δ_R^{++} and Δ_L triplet of states do not mix with other scalars, and their mass matrix is diagonal to begin with. The simplest is

$$m_{\Delta_R^{++}}^2 = v_R^2 \left(4\rho_2 + \frac{c_{2\beta}}{c_\beta^4} \alpha_3 \epsilon^2 \right) , \qquad (50)$$

and the Δ_L triplet masses are given by

$$m_{\Delta_L^0}^2 = v_R^2 \left[\rho_3 - 2\rho_1 + \epsilon^2 \alpha_3 \, s_{2\beta} t_{2\beta} \, s_\alpha^2 \right] = m_{\chi_L^0}^2 \,, \qquad (51)$$

$$m_{\Delta_{L}^{+}}^{2} = v_{R}^{2} \bigg[\rho_{3} - 2\rho_{1} + \epsilon^{2} \alpha_{3} \left(s_{2\beta} t_{2\beta} s_{\alpha}^{2} + \frac{c_{2\beta}}{2} \right) \bigg], \quad (52)$$

$$m_{\Delta_{L}^{++}}^{2} = v_{R}^{2} \bigg[\rho_{3} - 2\rho_{1} + \epsilon^{2} \alpha_{3} \left(s_{2\beta} t_{2\beta} s_{\alpha}^{2} + c_{2\beta} \right) \bigg], \quad (53)$$

where clearly the four states within the Δ_L triplet are degenerate up to ϵ^2 corrections. These relations can be readily solved for e.g. ρ_2 and ρ_3 in terms of two masses, $m_{\Delta_R^{++}}$, m_{Δ_L} , as we do below. The remaining three masses within

the Δ_L triplet are then predicted and follow the known sum rule [99]

$$m_{\Delta_L^{++}}^2 - m_{\Delta_L^{+}}^2 = m_{\Delta_L^{+}}^2 - m_{\Delta_L^{0}}^2 = v^2 \alpha_3 \frac{c_{2\beta}}{2}.$$
 (54)

2.2.2 Singly charged states (H^+ and the Goldstones of $W_{L,R}$)

From the three singly charged states Φ_1^+ , Φ_2^+ , Δ_R^+ , one finds two Goldstones relative to W_L^+ and W_R^+ , and only one massive singly charged state, with mass

$$m_{H^+}^2 = v_R^2 \frac{\alpha_3}{c_{2\beta}} \left(1 + \frac{c_{2\beta}^2}{2} \epsilon^2 \right).$$
 (55)

The unitary rotation from the unphysical to the physical basis was described in (28) and (32).

2.2.3 Neutral states (h, Δ , H, A plus the two Goldstones of Z_L , Z_R)-restriction $\lambda_3 = 2\lambda_2$

After the t_{β} demixing in (36), the neutral goldstone bosons are decoupled from the neutral massive states { $\Re \varphi_{10}$, $\Re \Delta_R$, $\Re \varphi_{20}$, $\Im \varphi_{20}$ }. Their mass matrix is then diagonalized to yield the mass eigenstates {h, Δ , H, A}.

It is convenient to restrict to $\lambda_3 = 2\lambda_2$, which removes a tiny and phenomenologically irrelevant CP-violating mixing between the *H* and *A* states, and allows for almost exact diagonalization. The 4 × 4 squared mass matrix becomes, in units of v_R^2 ,

The diagonalization proceeds in two steps.

• We first notice that the 2-1 entry, driving the mixing θ of the SM higgs with the unphysical Δ (i.e. $h - \Re \Delta_R$) is constrained by exotic and invisible Higgs decay searches to be phenomenologically small: $|\sin \theta| < 20\%$ [119]. The two lower entries are also small, $\sim \epsilon^2$. Thus, all three entries can be cleaned with rotations at leading order. This isolates the Higgs eigenvalue, which can be expressed in terms of the $h - \Re \Delta_R$ mixing angle θ (reported below)

$$m_{h}^{2} = v^{2} \left(4\lambda_{1} + \frac{64\lambda_{2}t^{2}c_{\alpha}^{2}}{\left(t^{2} + 1\right)^{2}} + \frac{16\lambda_{4}tc_{\alpha}}{t^{2} + 1} - Y\tilde{\theta}^{2} \right), \quad (57)$$

where we use the shorthand $\tilde{\theta} \equiv \theta/\epsilon \lesssim O(1)$.

Then, remarkably thanks to the λ₃ = 2λ₂ restriction, the remaining 3 × 3 block is diagonalized exactly by two rotations, in the 3 – 4 and 2 – 3 planes, instead of three. This isolates the last three mass eigenvalues, in terms of the 3 – 4 rotation.

The simplest eigenvalue belongs to the pseudoscalar, which we used to compactify the notation

$$m_A^2 = v_R^2 X, \qquad X \equiv \frac{1+t^2}{1-t^2} \alpha_3.$$
 (58)

It is clearly degenerate with H^+ , see (55), up to ϵ^2 terms.

$$\begin{pmatrix} 4\epsilon^{2} \left(\lambda_{1} + \frac{4tc_{\alpha}(\lambda_{4}(t^{2}+1)+4\lambda_{2}tc_{\alpha})}{(t^{2}+1)^{2}}\right) 2\epsilon \left(\alpha_{1} - \frac{tX(t^{3}+t-2\eta_{2}c_{\alpha})}{(t^{2}+1)^{2}}\right) \frac{4\epsilon^{2}(t^{2}c_{2\alpha}-1)(\lambda_{4}(t^{2}+1)+8\lambda_{2}tc_{\alpha})}{(t^{2}+1)^{2}} \frac{4t^{2}\epsilon^{2}s_{2\alpha}(\lambda_{4}(t^{2}+1)+8\lambda_{2}tc_{\alpha})}{(t^{2}+1)^{2}} \\ 2\epsilon \left(\alpha_{1} - \frac{tX(t^{3}+t-2\eta_{2}c_{\alpha})}{(t^{2}+1)^{2}}\right) Y \frac{2X\epsilon(t^{2}c_{2\alpha}-1)\eta_{2}}{(t^{2}+1)^{2}} \frac{2t^{2}X\epsilon_{s_{2\alpha}\eta_{2}}}{(t^{2}+1)^{2}} \\ \frac{4\epsilon^{2}(t^{2}c_{2\alpha}-1)(\lambda_{4}(t^{2}+1)+8\lambda_{2}tc_{\alpha})}{(t^{2}+1)^{2}} \frac{2X\epsilon(t^{2}c_{2\alpha}-1)\eta_{2}}{(t^{2}+1)^{2}} X + \frac{16\lambda_{2}\epsilon^{2}(t^{2}c_{2\alpha}-1)}{(t^{2}+1)^{2}} \frac{16\lambda_{2}t^{2}\epsilon^{2}s_{2\alpha}(t^{2}c_{2\alpha}-1)}{(t^{2}+1)^{2}} \\ \frac{4t^{2}\epsilon^{2}s_{2\alpha}(\lambda_{4}(t^{2}+1)+8\lambda_{2}tc_{\alpha})}{(t^{2}+1)^{2}} \frac{2t^{2}X\epsilon_{s_{2\alpha}\eta_{2}}}{(t^{2}+1)^{2}} \frac{16\lambda_{2}t^{2}\epsilon^{2}s_{2\alpha}(t^{2}c_{2\alpha}-1)}{(t^{2}+1)^{2}} X + \frac{16\lambda_{2}t^{4}\epsilon^{2}s_{2\alpha}^{2}}{(t^{2}+1)^{2}} \end{pmatrix}$$
(56)

where for compactness we defined $X \equiv \frac{1+t^2}{1-t^2}\alpha_3$, $Y \equiv 4\rho_1$, $t \equiv t_\beta$, and $\eta_2 \equiv t \sin(\alpha + \delta_2) / \sin \delta_2$.

The remaining states, after this and the SM Higgs in (57), are the neutral triplet and the heavy scalar

 $\theta \equiv$

$$m_{\Delta}^{2} = v_{R}^{2} \left[Y + \sec(2\eta) \left[(Y - X)s_{\eta}^{2} + \epsilon^{2} \left(Y \tilde{\theta}^{2} c_{\eta}^{2} - \frac{16\lambda_{2} \left(t^{4} - 2c_{2\alpha} t^{2} + 1 \right)}{\left(t^{2} + 1 \right)^{2}} s_{\eta}^{2} \right) \right] \right],$$
(59)

$$m_{H}^{2} = v_{R}^{2} \left[X - \sec(2\eta) \left[(Y - X)s_{\eta}^{2} + \epsilon^{2} \left(Y \tilde{\theta}^{2} s_{\eta}^{2} - \frac{16\lambda_{2} \left(t^{4} - 2c_{2\alpha} t^{2} + 1 \right)}{\left(t^{2} + 1 \right)^{2}} c_{\eta}^{2} \right) \right] \right].$$
(60)

Here we list the subsequent orthogonal rotations, complementing the neutral GB rotations defined in (39):

$$\epsilon \,\tilde{\theta} \equiv \theta_{21} = \epsilon \frac{2\,\alpha_1 - 2\,X\,t\left(\frac{t}{1+t^2} - \frac{2\,\eta_2\,c_\alpha}{(t^2+1)^2}\right)}{Y - 4\,\epsilon^2\left(\lambda_1 + 16\lambda_2\frac{t^2\,c_\alpha^2}{(t^2+1)^2} + 4\lambda_4\frac{t\,c_\alpha}{t^2+1}\right)} \tag{61}$$

$$\phi \equiv \epsilon^2 \tilde{\phi} \equiv \theta_{31} = \epsilon^2 \frac{(t^2 c_{2\alpha} - 1)}{(t^2 + 1)^2} \left[\frac{32 t c_\alpha \lambda_2 + 4 \lambda_4 (t^2 + 1)}{X} - 2 \tilde{\theta} \eta_2 \right], \tag{62}$$

$$\theta_{41} = \epsilon^2 \frac{t^2 s_{2\alpha}}{\left(t^2 + 1\right)^2} \left[\frac{32 t c_\alpha \lambda_2 + 4 \lambda_4 (t^2 + 1)}{X} - 2 \tilde{\theta} \eta_2 \right],\tag{63}$$

$$\theta_{34} = \tan^{-1} \left(\frac{t^2 s_{2\alpha}}{t^2 c_{2\alpha} - 1} \right),\tag{64}$$

$$\eta \equiv \theta_{23} = -\frac{1}{2} \tan^{-1} \left[\frac{4 X \epsilon \sqrt{t^4 - 2c_{2\alpha}t^2 + 1} \eta_2}{\left(t^2 + 1\right)^2 \left(Y \tilde{\theta}^2 \epsilon^2 - \frac{16(t^4 - 2c_{2\alpha}t^2 + 1)\lambda_2 \epsilon^2}{\left(t^2 + 1\right)^2} - X + Y\right)} \right].$$
(65)

The combined rotation for neutral scalars that relates the unphysical to the mass basis is thus

$$\begin{pmatrix} \Re\varphi_{10}\\ \Re\Delta_R\\ \Re\varphi_{20}\\ \Im\varphi_{20} \end{pmatrix} = O_N \begin{pmatrix} h\\ \Delta\\ H\\ A \end{pmatrix}, \tag{66}$$

with $O_N = O_{21}O_{31}O_{41}O_{34}O_{23}$, where each

$$O_{ij} = \begin{pmatrix} c_{\theta_{ij}} & -s_{\theta_{ij}} \\ s_{\theta_{ij}} & c_{\theta_{ij}} \end{pmatrix}$$
(67)

acts in the (i, j) subspace, with angle θ_{ij} from (61)–(65).

The first angle has to be small by Higgs decay constraints. This implies e.g. a small α_1 . In case one considers a light m_{Δ} to the EW scale or below, $Y \leq \epsilon^2$, the ϵ^2 correction in the denominator of Eq. (61) becomes important. The numerator has clearly to be even smaller.

The last angle η instead can be $\mathcal{O}(1)$, provided X - Y are degenerate to $\mathcal{O}(\epsilon)$, which corresponds to a H - A degeneracy of the order of 500 GeV or less.

In the next section all previous expressions for mixings and for masses are inverted so that couplings are expressed in terms of chosen physical quantities. In particular, with the above convention, the interesting angles can be defined as

 $h \text{ part of } \Re \Delta_R : \ \theta \equiv \theta_{21} \simeq -(O_N)_{2,1},$ (68)

H part of $\Re \Delta_R$: $\eta \equiv \theta_{23} = \arcsin[(O_N)_{2,3}/c_{\theta}],$ (69)

h part of
$$\Re \phi_{20}$$
: $\phi \equiv \theta_{31} \simeq -(O_N)_{3,1}$, (70)

which we will use as physical inputs. Notice that if Δ , H, A are left around the scale of v_R , one has the following scalings $\eta \sim 1, \theta \sim \epsilon, \phi \sim \epsilon^2$.

It is nice to check the case of vanishing $t_{\beta} \rightarrow 0$: in this limit, λ_2 governs the *A*-*H* mass splitting, λ_4 governs the ϕ mixing and α_1 sets the θ mixing. On the other hand, the CP-violating (*A*-*h* and *A*-*H*) mixings θ_{41} and θ_{34} vanish as expected. The mixing η is governed by η_2 , which in this limit can remain nonzero, see below.

Finally, comparing the heavy scalar masses in Eqs. (58) and (55), we find another sum rule like the one in (54):

$$m_{H^+}^2 - m_A^2 = v^2 \alpha_3 \frac{c_{2\beta}}{2} \sim O(150 \,\text{GeV})^2 \,.$$
 (71)

It implies, recalling that A, H have to be as heavy as 20 TeV[18-26], and also that the splitting with H^+ becomes rather tiny, 1–10 GeV.

2.3 Inversion: couplings in terms of physical masses and mixings

Equations (57–58) together with (61), (62) and (65) can be solved for the relevant couplings of the potential, in terms of four masses $(m_h^2, m_{\Delta}^2, m_H^2, m_A^2)$ and the three mixings θ , η and ϕ . Collecting all the previous expressions one finds:

$$\frac{1+t^2}{1-t^2}\alpha_3 \equiv X \to \frac{m_A^2}{v_R^2},\tag{72}$$

$$4\rho_1 \equiv Y \to \frac{m_{\Delta}^2 + (m_H^2 - m_{\Delta}^2) s_{\eta}^2}{v_R^2},$$
(73)

$$\rho_2 \to \frac{m_{\Delta_R^{++}}^2}{4v_R^2} - \frac{\epsilon^2 X \left(t^2 - 1\right)^2}{4 \left(t^2 + 1\right)^2},\tag{74}$$

$$\rho_3 \to \frac{m_{\Delta_L}^2}{v_R^2} + \frac{Y}{2} - \frac{4\epsilon^2 X t^2 s_{\alpha}^2}{\left(t^2 + 1\right)^2},$$
(75)

$$\lambda_1 \to \frac{m_h^2}{4v^2} + \frac{Y}{4}\tilde{\theta}^2 - 4tc_\alpha \left(\frac{\lambda_4}{t^2+1} + \frac{4tc_\alpha\lambda_2}{\left(t^2+1\right)^2}\right), \quad (76)$$

$$\lambda_2 \to \frac{m_H^2 - m_A^2 - (m_H^2 - m_\Delta^2) s_\eta^2}{16 v^2} \frac{(t^2 + 1)^2}{(t^4 - 2c_{2\alpha}t^2 + 1)},$$
(77)

$$\lambda_3 \to 2\lambda_2 \,, \tag{78}$$

$$\lambda_4 \to \tilde{\phi} \frac{X(t^2+1)}{4(t^2 c_{2\alpha} - 1)} + \tilde{\theta} \frac{X\eta_2}{2(t^2+1)} - 8 \frac{tc_{\alpha}\lambda_2}{(t^2+1)}$$
(79)

$$\eta_2 \to \frac{(m_H^2 - m_{\Delta}^2) (t^2 + 1)^2 \sin(2\eta)}{4m_A^2 \epsilon \sqrt{t^4 - 2c_{2\alpha}t^2 + 1}}$$
(80)

$$\delta_2 \to \tan^{-1}\left(\frac{ts_{\alpha}}{\eta_2 - tc_{\alpha}}\right),$$
(81)

$$\alpha_1 \to \frac{Y\tilde{\theta}}{2} + t\left(\frac{t\,X}{t^2 + 1} + \frac{(Y - X)c_{\alpha}t_{2\eta}}{2\,\epsilon\sqrt{t^4 - 2c_{2\alpha}t^2 + 1}}\right),\tag{82}$$

$$\alpha_2 \to \frac{X}{2} \frac{(\eta_2 - tc_\alpha)}{(t^2 + 1)} \sqrt{\frac{t^2 s_\alpha^2}{(\eta_2 - tc_\alpha)^2}} + 1.$$
(83)

These expressions have been written by ensuring that no singularity develops for $t \to 0$ and/or $\alpha \to 0$.

On the other hand, a few comments are in order about the perturbativity of the couplings obtained, while referring to [90] for a thorough study of perturbativity limits, including loop corrections. Raising the scalar masses too much beyond v_R , like for instance m_A in (72), produces couplings that enter in a non-perturbative regime. But also, a few couplings are sensitive to scalar mass splittings in relation to mixing angles. In particular, λ_2 becomes large in case $|m_H^2 - m_A^2| > 16v^2$, or in case $|m_H^2 - m_\Delta^2|$ is similarly large while demanding large η mixing. Similar considerations apply to α_1 , if $(m_H^2 - m_\Delta^2)/\epsilon$ becomes too large. Next, it is obviously difficult to have a

large $h - \Delta$ mixing θ near its phenomenological limit 0.2, if m_{Δ} becomes too large. One runs towards non-perturbativity in $\alpha_1 \sim \theta Y/2\epsilon \sim m_{\Delta}/10 v$. Namely, one expectedly can not have a large h- Δ mixing if Δ is heavier than circa 1–2 TeV.

For all these reasons, a check on couplings shall be performed while varying input parameters, to avoid unphysical situations.

For testing purposes, a function that performs the inversion along with this check is provided along with the model package. The inversion above together with the other coupling constants solved previously, is implemented as a Mathematica function LRSMEVAL [expr, inputs], and provided as additional material [120, 121]. Here, inputs is a set of rules for the relevant needed input parameters, where input masses are to be expressed in TeV (or as a function of other known masses). For instance,

$$\begin{aligned} \text{inputs} &= \left\{ \text{MWR} \rightarrow 10, \text{ mH} \rightarrow 1.1 \text{ MWR}, \\ \text{mA} \rightarrow 1.1 \text{ MWR} + 0.1 \text{ MW}, \text{m}\Delta \rightarrow \text{MWR}, \\ \text{m}\Delta\text{Rpp} \rightarrow 1.5 \text{ MWR}, \text{m}\Delta\text{L} \rightarrow 2.5 \text{ MWR}, \\ \theta \rightarrow 0.05, \eta \rightarrow 0.01, \phi \rightarrow 0.0001, \\ \text{t} \rightarrow 0.01, \alpha \rightarrow 0.1, \rho 4 \rightarrow 0.5 \right\}. \end{aligned}$$

$$\end{aligned}$$

$$\end{aligned}$$

$$\end{aligned}$$

$$\end{aligned}$$

$$\end{aligned}$$

The function LRSMEVAL evaluates the given expr, function of couplings and masses, for the given physical inputs, while checking for perturbativity. It prints out a warning for each scalar coupling that may turn out to be larger than 3. The function was checked against exact numerical diagonalization of the relevant mass matrices. A further implementation in **Python** is also provided, see below.

2.4 Fermion sector

The LRSM fermions transform (in the interaction basis) under $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$ as

$$Q'_{Li} = \begin{pmatrix} u'_L \\ d'_L \end{pmatrix}_i \sim \left(\mathbf{2}, \mathbf{1}, \frac{\mathbf{1}}{\mathbf{3}}\right),\tag{85}$$

$$Q'_{Ri} = \begin{pmatrix} u'_R \\ d'_R \end{pmatrix}_i \sim \left(1, 2, \frac{1}{3}\right), \tag{86}$$

$$L'_{Li} = \begin{pmatrix} \nu'_L \\ \ell'_L \end{pmatrix}_i \sim (\mathbf{2}, \mathbf{1}, -\mathbf{1}), \qquad (87)$$

$$L'_{Ri} = \begin{pmatrix} \nu'_R \\ \ell'_R \end{pmatrix}_i \sim (\mathbf{1}, \mathbf{2}, -\mathbf{1}), \qquad (88)$$

and gauge invariant Yukawa terms are written as

$$\mathcal{L}_{Y}^{q} = \bar{Q}_{L}^{\prime} \left(Y_{q} \phi + \tilde{Y}_{q} \tilde{\phi} \right) Q_{R}^{\prime} + \text{H.c.}, \qquad (89)$$

$$\mathcal{L}_{Y}^{\ell} = \bar{L}_{L}^{\prime} \left(Y_{\ell} \phi + \tilde{Y}_{\ell} \tilde{\phi} \right) L_{R}^{\prime} + \bar{L}_{L}^{\prime c} i \sigma_{2} \Delta_{L} Y_{L}^{M} L_{L}^{\prime} + \bar{L}_{R}^{\prime c} i \sigma_{2} \Delta_{R} Y_{R}^{M} L_{R}^{\prime} + \text{H.c..}$$

$$(90)$$

All fermions thus get their masses exclusively from spontaneous symmetry breaking.

2.4.1 Quarks: Yukawa terms, masses and mixings

After the bi-doublet acquires its vevs, see Eq. (4), the mass matrices of the quarks can be written as

$$M_u = Y_q \, v_1 - \tilde{Y}_q \, e^{-i\alpha} v_2 \,, \tag{91}$$

$$M_d = -Y_q \, e^{i\alpha} v_2 + \tilde{Y}_q \, v_1 \,. \tag{92}$$

In general, these are arbitrary complex 3×3 mass matrices that can be diagonalised with a bi-unitary transformation

$$M_{u} = U_{uL} m_{u} U_{uR}^{\dagger}, \qquad M_{d} = U_{dL} m_{d} U_{dR}^{\dagger}, \qquad (93)$$

where m_u and m_d are diagonal with positive eigenvalues and $U_x U_x^{\dagger} = 1$. What enters into the physical charged current interactions are the usual CKM matrix $V_L^{\text{CKM}} \equiv U_{uL}^{\dagger} U_{dL}$ and its right-handed analog $V_R = U_{uR}^{\dagger} U_{dR}$, which in principle has different angles and five extra phases. The additional CP phases can be extracted as

$$V_R = K_u V_R^{\rm CKM} K_d \,, \tag{94}$$

where V_R^{CKM} is the right-handed analog of the CKM matrix and $K_{u,d}$ are diagonal matrices of complex phases

$$K_u = \operatorname{diag}\left(e^{i\theta_u}, e^{i\theta_c}, e^{i\theta_t}\right), \qquad (95)$$

$$K_d = \operatorname{diag}\left(e^{i\theta_d}, e^{i\theta_s}, e^{i\theta_b}\right), \qquad (96)$$

out of which only five are physical. The quark Yukawa couplings that enter in the Lagrangian can be re-written in terms of the physical masses and mixings:

$$Y_q = \frac{1}{v_1^2 - v_2^2} \left(M_u \, v_1 + M_d \, e^{-i\alpha} v_2 \right) \,, \tag{97}$$

$$\tilde{Y}_q = -\frac{1}{v_1^2 - v_2^2} \left(M_d \, v_1 + M_u \, e^{i\alpha} \, v_2 \right) \,. \tag{98}$$

Already from the third generation, one understands from here that the region of $t_{\beta} \leq 0.5$ leads to large non-perturbative couplings and should be excluded, see Eq. (2.14) in [77] and the Appendix A of [20] for details.

2.4.2 Leptons: Yukawa terms, Dirac and Majorana masses and mixings

In the lepton sector, after the bi-doublet and the scalar triplets acquire their respective vevs, we have the usual Dirac mass for the charged leptons and a type I + II seesaw mechanism in the neutrino sector. The charged lepton masses and the Dirac-Yukawa term between the left- and right-handed neutrinos can be cast as

$$M_D = Y_\ell v_1 - \tilde{Y}_\ell e^{-i\alpha} v_2 \,, \tag{99}$$

$$M_{\ell} = -Y_{\ell} e^{i\alpha} v_2 + \tilde{Y}_{\ell} v_1 \,. \tag{100}$$

In the same way, we can also rewrite the lepton Yukawa couplings in terms of the charged lepton masses and the Dirac mass term for the neutrinos

$$Y_{\ell} = \frac{1}{v_1^2 - v_2^2} \left(M_D v_1 + M_{\ell} e^{-i\alpha} v_2 \right),$$

$$\tilde{Y}_{\ell} = -\frac{1}{v_1^2 - v_2^2} \left(M_{\ell} v_1 + M_D e^{i\alpha} v_2 \right).$$
(101)

Similarly to the quark masses, the charged lepton mass term can be diagonalised via a bi-unitary transformation

$$M_{\ell} = U_{\ell L} \, m_{\ell} \, U_{\ell R}^{\dagger} \,. \tag{102}$$

In the neutrino sector on the other hand, one has to write the mass-matrix as

$$\mathcal{L}_{\nu\text{-mass}} = -\frac{1}{2} \bar{n}'_L M_n n'^c_L$$

= $-\frac{1}{2} \left(\bar{\nu}'_L \ \bar{\nu}'^c_R \right) \begin{pmatrix} M_L \ M_D \\ M^T_D \ M_R \end{pmatrix} \begin{pmatrix} \nu'^c_L \\ \nu'_R \end{pmatrix} + \text{H.c.},$ (103)

in which $n'_L = (\bar{v}'_L \ \bar{v}'_R)$ and $M_L = v_L Y^M_L$ and $M_R = v_R Y^M_R$. This 6×6 Majorana mass matrix is by construction (complex) symmetric and can be diagonalised via an Autonne-Takagi factorisation.

The full mass-matrix can be perturbatively (block-) diagonalised with a unitary rotation

$$\tilde{W}^T M_n \tilde{W} = \begin{pmatrix} m_{\text{light}} & 0\\ 0 & m_{\text{heavy}} \end{pmatrix},$$
(104)

in which the rotation matrix can be defined as

$$\tilde{W} = \begin{pmatrix} \sqrt{\mathbb{1} - BB^{\dagger}} & B \\ -B^{\dagger} & \sqrt{\mathbb{1} - B^{\dagger}B} \end{pmatrix}.$$
(105)

The light and heavy mass eigenstates are up to leading order in M_R^{-1} given by

$$M_{\nu} \simeq M_L - M_D M_R^{-1} M_D^T, \qquad M_N \simeq M_R.$$
 (106)

The light and heavy mass matrices are then diagonalised via unitary rotations as

$$V_{\nu}^{T} M_{\nu} V_{\nu} = \text{diag}(m_{\nu_{1}}, m_{\nu_{2}}, m_{\nu_{3}}), \qquad (107)$$

$$V_N^{\mathsf{T}} M_N V_N^* = \text{diag}(m_{N_1}, m_{N_2}, m_{N_3}), \qquad (108)$$

in which the light states are identified with mostly active neutrinos. The full rotation matrix can be written as

$$W = \tilde{W} \begin{pmatrix} V_{\nu} & 0\\ 0 & V_N^* \end{pmatrix}.$$
(109)

We can further expand the unitary matrix W (and the matrix B) in inverse powers of M_R , resulting in

$$\sqrt{\mathbb{1} - BB^{\dagger}} \simeq \mathbb{1} - \frac{1}{2}B_1B_1^{\dagger} - \frac{1}{2}\left(B_1B_2^{\dagger} + B_2B_1^{\dagger}\right),$$
 (110)

which is valid up to $\mathcal{O}(M_R^{-4})$, with

$$B_1 = M_D^{\dagger} M_R^{-1\dagger}, \qquad B_2 = M_L^{\dagger} M_D^T M_R^{-1T} M_R^{-1\dagger}, \qquad (111)$$

see [122] for further details of the perturbative expansion. Up to leading order in M_R^{-1} , the full neutrino mixing matrix is then given by

$$W = \begin{pmatrix} \sqrt{1 - BB^{\dagger}V_{\nu}} & BV_{N}^{*} \\ -B^{\dagger}V_{\nu} & \sqrt{1 - B^{\dagger}B}V_{N}^{*} \end{pmatrix}$$

$$\simeq \begin{pmatrix} V_{\nu} & B_{1}V_{N}^{*} \\ -B_{1}^{\dagger}V_{\nu} & V_{N}^{*} \end{pmatrix}.$$
 (112)

The higher order corrections to the blocks of \tilde{W} become numerically relevant for the "off-diagonal" blocks, i.e. for the "light-heavy" mixing terms. With the neutrino and charged lepton mixing matrices we can now write the charged current interactions in the physical basis

$$\mathcal{L}_{cc}^{\ell} = \frac{g_L}{\sqrt{2}} \bar{\ell}_L \gamma^{\mu} \mathcal{U}_L n_L W_L^{\mu} + \frac{g_R}{\sqrt{2}} \bar{\ell}_R \gamma^{\mu} \mathcal{U}_R n_R W_R^{\mu}, \quad (113)$$

where the physical neutrino mass eigenstates are $n = (v_1, v_2, v_3, N_1, N_2, N_3)^T$, and where we defined the semiunitary 3×6 matrices

$$(\mathcal{U}_L)_{\alpha i} = \sum_{k=1}^3 (V_{\ell L}^{\dagger})_{\alpha k} W_{ki} , \qquad (114)$$

$$(\mathcal{U}_R)_{\alpha i} = \sum_{k=1}^{3} (V_{\ell L}^{\dagger})_{\alpha k} W_{(k+3)i}.$$
(115)

The first 3×3 block of U_L can be identified as the "wouldbe" PMNS matrix (which is in general no longer exactly unitary) and the second 3×3 block of U_R as its right-handed analog that can be directly accessed by experiments sensitive to right-handed currents.

Parametrising M_D in the case of C. In the case of C-invariance as the manifest LR symmetry, the scalars transform as $\phi \leftrightarrow \phi^T$ and $\Delta_L \leftrightarrow \Delta_R^*$, leading to convenient relations of the Yukawa couplings

$$Y_{\ell} = Y_{\ell}^{T}, \qquad \tilde{Y}_{\ell} = \tilde{Y}_{\ell}^{T}, \qquad Y_{L}^{M} = Y_{R}^{M}, \qquad (116)$$

and consequently for the mass matrices

$$M_D = M_D^T, \qquad \qquad M_L = \frac{v_L}{v_R} M_R. \qquad (117)$$

This allows us to invert Eq. (106) in order to obtain an expression of M_D , fully determined by physical parameters, in contrast to the additional freedom present in the general case. The Dirac mass matrix is then given by

$$M_D = M_N \sqrt{\frac{v_L}{v_R} \mathbb{1} - M_N^{-1} M_\nu} \,. \tag{118}$$

The square root of a matrix is in general quite a complicated expression, however it can be simplified to a relatively simple form

$$\sqrt{A} = c_0 \,\mathbb{1} + c_1 \,A + c_2 \,A.A,\tag{119}$$

which holds for any (also non-symmetric) complex 3×3 matrix *A*. The c_i coefficients are calculated in terms of the three invariants of *A* and are given in appendix A below.¹

2.5 Ghost sector

To complete the Faddeev–Popov gauge fixing procedure we have to deal with the ghosts, as in the case of the SM [118,124, 125]. In the LRSM we are dealing with an extended gauge group and gauge bosons that mix, therefore the procedure is slightly more involved. We need to introduce a set of ghost fields c_i that are associated with the gauge parameters α_i in the following way

$$c_i = \left\{ c_L^+, c_L^-, c_R^+, c_R^-, c_{Z_L}, c_{Z_R}, c_A \right\},$$
(120)

$$\alpha_i = \left\{ \alpha_L^+, \alpha_L^-, \alpha_R^+, \alpha_R^-, \alpha_{Z_L}, \alpha_{Z_R}, \alpha_A \right\}.$$
(121)

¹ In the case of no manifest left-right symmetry, M_D may still be parametrised via an analogue of the Casas-Ibarra parametrisation, uncovering the additional free parameters (see e.g. [123]). In the case of \mathcal{P} as the manifest left-right symmetry, additional relations and restrictions apply and the parametrisation of the lepton sector becomes significantly more complicated (see e.g. [80]).

These ghosts are written in the mass basis and enter the ghost Lagrangian in the usual way

$$\mathcal{L}_{\text{ghost}} = \sum_{i,j} \overline{c}_i \frac{\partial \left(\delta F_i\right)}{\partial \alpha_j} c_j , \qquad (122)$$

$$F_{i} = \left\{ F_{L}^{+}, F_{L}^{-}, F_{R}^{+}, F_{R}^{-}, F_{Z_{L}}, F_{Z_{R}}, F_{A} \right\},$$
(123)

where the sums run over the gauge parameters in the mass basis. The variations of the Fs are given by

$$\delta F_i^+ = \partial \delta W_i^+ + i \xi_{W_i} M_{W_i} \delta \varphi_i^+ , \qquad (124)$$

$$\delta F_{Z_i} = \partial \delta Z_i + \xi_{Z_i} M_{Z_i} \delta \varphi_i^0, \qquad (125)$$

$$\delta F_A = \partial \delta A \,, \tag{126}$$

and depend on the variations of the wbGs $\delta \varphi$ and the gauge fields δV . We get those by considering an infinitesimal expansion of the gauge transformations

$$\phi \to U\phi = \phi + \delta\phi \,, \tag{127}$$

$$D\phi \to U(D\phi) = D\phi + \delta(D\phi)$$
. (128)

The first term gives us the variations of the wbG and Higgs fields with respect to α_i . The second term gives the variation of the gauge fields, needed to restore gauge invariance of covariant derivatives. The bi-doublet and the triplet transform in the following way

$$\phi \to U_L \phi U_R^{\dagger}, \qquad \Delta_R \to U_R \Delta_R U_R^{\dagger} U_{B-L}, \quad (129)$$

$$U_{L,R} = e^{ig\alpha^a \frac{\sigma^a}{2}}, \qquad U_{B-L} = e^{ig' \frac{B-L}{2}\alpha^7}.$$
 (130)

The σ matrices in $U_{L,R}$ run over their respective group indices, setting a = 1, 2, 3 (a = 4, 5, 6) for the $SU(2)_{L(R)}$ group parameters in the gauge basis; α^7 parametrizes the $U(1)_{B-L}$ Abelian part. The field variations from (2) are

$$\delta\phi = \begin{pmatrix} \delta\phi_1^{0*} \ \delta\phi_2^+ \\ \delta\phi_1^- \ \delta\phi_2^0 \end{pmatrix}, \qquad \delta\Delta_R = \begin{pmatrix} \frac{\delta\Delta_R^+}{\sqrt{2}} & 0 \\ \delta\Delta_R^0 & -\frac{\delta\Delta_R^+}{\sqrt{2}} \end{pmatrix}, \qquad (131)$$

where we focus only on the variations of the wbG, i.e. $\delta \varphi_{L,R}^{\pm}$, $\delta \varphi_{L,R}^{0}$. We can ignore variations of all the other fields $(\delta \Delta_L \text{ multiplet}, \delta \Delta_R^{++}, \delta H^+, \text{ as well as the variations of the neutral scalars) which do not enter in <math>F_i$. Expanding the gauge transformations for small α^i and multiplying the generator and field matrices, we get the following variations for the bi-doublet and the triplet

$$\delta\phi = i\frac{g}{2}\left(\Sigma_L\phi - \phi\Sigma_R\right),\tag{132}$$

$$\delta \Delta_R = i \frac{g}{2} \left(\Sigma_R \Delta_R - \Delta_R \Sigma_R \right) + i g' \Delta_R \alpha^7 \,. \tag{133}$$

with

$$\Sigma_L = \begin{pmatrix} \alpha^3 & \alpha^1 - i\alpha^2 \\ \alpha^1 + i\alpha^2 & -\alpha^3 \end{pmatrix},$$

$$\Sigma_R = \begin{pmatrix} \alpha^6 & \alpha^4 - i\alpha^5 \\ \alpha^4 + i\alpha^5 & -\alpha^6 \end{pmatrix}.$$
(134)

We now perform all the rotations of the fields using U_+ and O_{φ} to go to the mass basis. We then define the α_i and c_i as the gauge parameters and ghost fields in the same mass basis as the gauge bosons. Specifically, we use the following relations

$$\frac{\sqrt{2}\,\tilde{\alpha}_{L}^{\pm} = \alpha^{1} \mp i\alpha^{2}}{\sqrt{2}\,\tilde{\alpha}_{R}^{\pm} = \alpha^{4} \mp i\alpha^{5}}, \begin{pmatrix} \tilde{\alpha}_{L}^{+} \\ \tilde{\alpha}_{R}^{+} \end{pmatrix} = U_{W} \begin{pmatrix} \alpha_{L}^{+} \\ \alpha_{R}^{+} \end{pmatrix},$$
(135)

to reparametrize the 'charged' gauge parameters (and the same for ghosts when $c^i \rightarrow c_{L,R}^{\pm}$). Likewise, we utilize the neutral gauge rotations for the remaining

$$\begin{pmatrix} \alpha^3 \\ \alpha^6 \\ \alpha^7 \end{pmatrix} = O_Z \begin{pmatrix} \alpha_A \\ \alpha_{Z_L} \\ \alpha_{Z_R} \end{pmatrix}.$$
 (136)

To get the gauge variations δV_i in the mass basis, we proceed along similar lines. We start by deriving δV_i in the flavor basis by demanding that the covariant derivatives, defined in (9) and (10), transform like the corresponding fields ϕ and Δ_R

$$D\phi \to U_L \ (D\phi) \ U_R^{\dagger},$$
 (137)

$$D\Delta_R \to U_R \left(D\Delta_R \right) U_R^{\dagger} U_{B-L} \,.$$
 (138)

We use the same expansions in α of the gauge transformations $U_{L,R,B-L}$ as above and apply the same rotations U_W and O_Z for δW , δZ , as we did for the W and Z. With both $\delta \phi$ and δV , we now have all the δF_i written as linear functions of the α^i . It is thus trivial to take the derivatives $\partial \delta F_i / \partial \alpha^i$ over the group parameters and plug them into (122) to finalize the ghost Lagrangian.

3 Left-right asymmetric case

In the previous sections we assumed that parity is manifestly respected in the gauge sector by imposing $g_L = g_R$. Such an equality should hold at the scale of parity restoration, which may happen at a high scale, such that the two couplings $g_{L,R}$ run differently and $g_L \neq g_R$ at a low scale. Although the running leads to a quite small splitting of gauge couplings, other extensions of the LRSM may lead to different $g_{L,R}$ with no other appreciable variations. It is thus interesting to generalize the results to this possibility. Let us then set $g_L = g$, $g_R = \zeta g$ and $g' = rg_R = r\zeta g$. This change only affects the various covariant derivatives and does not enter in the potential, such that the expressions for scalar masses and mixing (including the Goldstones) remain the same.

Charged gauge bosons. The mass matrix for charged gauge bosons now becomes

$$M_{W_{LR}}^2 = \frac{g^2}{2} v_R^2 \begin{pmatrix} \epsilon^2 & e^{-i\alpha} \epsilon^2 s_{2\beta} \zeta \\ e^{i\alpha} \epsilon^2 s_{2\beta} \zeta & (2+\epsilon^2) \zeta^2 \end{pmatrix}.$$
 (139)

The light eigenvalue remains the same $M_{W_L} \simeq gv/\sqrt{2}$, but the heavier one gets multiplied with ζ , such that $M_{W_R} \simeq gv_R (1 + \epsilon^2/4) \zeta$. The mixing angle is also different and is given by

$$s_{\xi} \simeq \frac{\epsilon^2}{2\zeta} s_{2\beta} \simeq \frac{M_{W_L}^2}{M_{W_R}^2} s_{2\beta} \zeta . \tag{140}$$

Up to the relevant order $\mathcal{O}(\epsilon^2)$, the Goldstone rotations in Eq. (32) are not affected by a change in g_R , such that U_+ remains the same as in the parity conserving case. Of course the ζ dependence enters in the actual terms, but the solution for the rotation angle is ζ independent. The same holds for the gauge fixing F_i^{\pm} terms, once the ζ dependence is absorbed in M_{W_R} .

Neutral gauge bosons. With explicit parity breaking, the neutral gauge boson mass matrix is

$$M_{Z_{LR}}^{2} = \frac{g^{2}}{2} v_{R}^{2} \begin{pmatrix} \epsilon^{2} & -\epsilon^{2}\zeta & 0\\ -\epsilon^{2}\zeta & (4+\epsilon^{2})\zeta^{2} & -4r\zeta\\ 0 & -4r\zeta & 4r^{2} \end{pmatrix}.$$
 (141)

and the modified eigenvalues are

$$M_Z \simeq \frac{gv}{\sqrt{2}} \sqrt{1 + \frac{r^2 \zeta^2}{r^2 + \zeta^2}},$$
 (142)

$$M_{Z_{LR}} \simeq g v_R \sqrt{2(\zeta^2 + r^2)} \left[1 + \frac{\epsilon^2 \zeta^4}{8 \left(\zeta^2 + r^2\right)^2} \right].$$
(143)

The definition of the weak mixing angle also impacts the *r*-parameter

$$r = \frac{t_w}{\sqrt{\zeta^2 - t_w^2}} \xrightarrow{\zeta \to 1} \frac{s_w}{\sqrt{c_{2w}}},\tag{144}$$

and the gauge boson rotation matrix gets updated to

$$O_Z = \begin{pmatrix} s_w & -c_w & 0\\ \frac{s_w}{\zeta} & \frac{s_w t_w}{\zeta} & -\frac{\sqrt{\zeta^2 - t_w^2}}{\zeta}\\ \frac{c_w \sqrt{\zeta^2 - t_w^2}}{\zeta} & \frac{s_w \sqrt{\zeta^2 - t_w^2}}{\zeta} & \frac{t_w}{\zeta} \end{pmatrix}$$

Table 1 FeynRules files included in the mLRSM package

mLRSM.fr	Main file (full)
mLRSM_UG.fr	Main file (unitary gauge)
LRSM_all_parameters.fr	External inputs (w/o masses)
LRSM_potential.fr	Scalar Potential
LRSM_lepton_yukawa.fr	Lepton Yukawa definitions
LRSM_quark_yukawa.fr	Quark Yukawa definitions
LRSM_matrices.fr	Matrix utilities
LRSM_ONdef.fr	Scalar mixing matrix
LRSM_Massless_5f.rst	Restriction $m_{u,d,s,c,b,e,\mu} = 0$
LRSM_Massless_5f_nu.rst	Restriction $m_{u,d,s,c,b,e,\mu,\nu_i} = 0$

$$+ \frac{\epsilon^{2}}{4} \begin{pmatrix} 0 & 0 & \frac{\left(\zeta^{2} - t_{w}^{2}\right)^{\frac{3}{2}}}{\zeta^{4}} \\ 0 & -\frac{\left(\zeta^{2} - t_{w}^{2}\right)^{2}}{c_{w}\zeta^{5}} & -\frac{t_{w}^{2}\left(\zeta^{2} - t_{w}^{2}\right)^{\frac{3}{2}}}{\zeta^{5}} \\ 0 & \frac{s_{w}\left(\zeta^{2} - t_{w}^{2}\right)^{\frac{3}{2}}}{c_{w}^{2}\zeta^{5}} & -\frac{t_{w}\left(\zeta^{2} - t_{w}^{2}\right)^{2}}{\zeta^{5}} \end{pmatrix}.$$
(145)

As with the singly charged Goldstones, also the neutral ones are rotated with the same matrix like in the parity conserving case at the ϵ^2 order. In particular, the O_{φ} remains the same as it was in (38). All the other parts of the Lagrangian (fermions, scalars) remain the same, except for the ghosts. The Lagrangian terms of the ghost sector are updated with $g_L = g, g_R \rightarrow \zeta g$ and $g' \rightarrow r\zeta g$, and their rotations are ζ -dependent, being the same as those of the gauge bosons.

As a final notice, the argument inside the square roots in the last expressions should always be positive, corresponding to the physical constraint, stemming from the symmetry breaking relations,

$$\frac{g_R}{g_L} \equiv \zeta > t_w \simeq 0.535. \tag{146}$$

4 The FeynRules implementation

The model file is split for cleanliness into a main model file plus various auxiliary files, reported in Table 1. In particular, the file LRSM_all_parameters.fr contains all external parameter definitions, as well all the expressions for the internal parameters calculated from the inversion described in Sect. 2.3. The lepton mass and mixing computations and the extraction of the square root are included in the files LRSM_lepton_yukawa.fr and LRSM_matrices.fr, and similarly for the quark matrices. The quite lengthy expression of the neutral scalars mixings (66) are included in LRSM_ONdef.fr.

In the full model mLRSM.fr we implemented both the Feynman and the Unitary gauge (Table 2). Either option can be adopted (e.g. before UFO generation) by defining

the switch variable FeynmanGauge=True(False). For usual unitary gauge studies, a simplified model file without ghosts and wbGs is provided as mLRSM_UG.fr. On the other hand, Feynman-gauge models are highly convenient for CalcHEP/CompHEP [126], where computations can become 10–100 times faster.

In Table 3 we list all the particles with their chosen masses for the benchmark point. Widths are also reported, as calculated from the leading decay diagrams, with the chosen mass spectrum. In case one or more masses are changed, the relevant widths have clearly to be recalculated for consistency.

Table 4 lists the other external parameters, including initial benchmark values, with the LHA blocks and counters. They include the important dimensionless parameters like t_{β} , α and scalar field mixings, as well as the mixing matrices for quarks (CKM) and leptons (PMNS). Finally, the lightneutrino mass parameters are defined as inputs, including the unknown lightest-neutrino mass and the binary choice between normal and inverted hierarchy.

Inside the model file a number of unphysical fields are defined, in order to hold fields in gauge multiplets, flavor multiplets, collection of Majorana neutrino fields, collections of neutral or charged scalars. The triplet fields $\Delta_{L,R}$ are arranged in two index (2, $\overline{2}$) representations, and the covariant derivative definition in FeynRules were generalized accordingly for them to work on reducible representations.

For collider studies, the "light" (dynamic) quarks u, d, s, c, b and leptons e, μ , as well as the light neutrinos v_i can usually be considered massless for all kinematical purposes. Setting to zero also the Dirac neutrino couplings leads to a considerable reduction of the complexity of model parameters and interactions, especially welcome at NLO order. Thus, along with the full model file, we provide two restriction files for such purpose, see Table 1.

4.1 UFO, parallelization and speedup

The FR model file can be used to produce the UFO package by using FeynRules; in this work we used the version Feynrules-2.3.49. The complexity of the present model largely exceeds that of previous ones and requires the highest possible level of parallelization. This was not fully present in the current version of FeynRules, where a number of issues prevents optimal running in parallel. A patched version is then provided as Feynrules-2.3.49_fne. With this patched version the time required to generate the LO UFO decreases from circa 1 week to O(1) hour on 10 cores. Similar speedup patches had to be applied to the MoGRe package [127], which is also provided as MoGRe_v1.1_fne.m, as well as to the Feynarts interface within FR. These speedup patches are essential for the generation of the NLO version, even with no neutrino Dirac parameters, see next subsection.

Table 2 UFOs included in the mLRSM package, with various restrictions. Here, the first three lines list UFO models in Feynman gauge; the last three instead list models (-ug) where unitary gauge was enforced, stripping off ghosts and wbGs. While e, μ are always massless, we provide -4f and -5f massless flavor schemes, and -nu stands for further simplifications with zero light neutrino masses and Dirac couplings. On the other hand, -full stands for models where all fermions are massive. Finally, -loop stands for NLO, requiring massless light quarks. (*) Notice that unitary gauge NLO UFOs are not correctly supported by Madgraph at the moment

UFO @ LO	UFO @ NLO
mlrsm	mlrsm-loop
mlrsm-nu	mlrsm-nu-loop
mlrsm-full	
mlrsm-ug	mlrsm-ug-loop(*)
mlrsm-ug-nu	mlrsm-ug-nu-loop(*)
mlrsm-ug-full	

The Mathematica notebooks used to generate the UFOs as well as the patched packages are available along with the model files as supplementary material, and uploaded to [120, 121].

4.2 NLO and restrictions

UFO packages at NLO in the QCD coupling can be generated from the above model files, by feeding first the relevant (QCD) subset of the Lagrangian into MoGRe [127]. The MoGRe package calculates the needed one-loop QCD couterterms, which are then processed with FeynArts [128] and NLOCT [129] to generate the one loop amplitudes, and these are used into FeynRules to generate the UFO at NLO.

We generated UFOs for a number of relevant model restrictions and cases, as listed in Table 2. The first three lines list models in the Feynman gauge, and the last three lines the models in the unitary gauge, devoid of ghosts and wbGs. The UG models should be sufficient for collider event generation at LO, while at NLO only Feynman gauge should be used with Madgraph, because loop diagrams are not correctly evaluated in the unitary gauge, at the moment [130].

Also, at NLO the u, d, s, c, b quarks are always restricted to be massless, which is required for numerical stability of the renormalization procedure. In case the light neutrino masses, as well as the Dirac couplings, are not needed, one may use the mlrsm-nu-loop model, where they were set to zero, retaining only masses for heavy N_i . The mlrsm-loop NLO model permits otherwise nonzero neutrino masses, parametrized by mixing and oscillation parameters, as well as lightest neutrino mass scale MNU0 and neutrino hierarchy NUH. We recall that in the model neutrino masses drive the Dirac mass matrix and Yukawa couplings, via Eqs. (118)

Table 3 Physical particle fields in the LRSM, with their masses, widths, and PDG IDs

Name	Spin	Self-Conj.	FR name	FR mass	Value [GeV]	FR width	Value [GeV]	PDG ID
G	1	Yes	g	MG	0	0		21
Α	1	Yes	A	MA	0	0		22
Z_L	1	Yes	Z	MZ	91.19	WZ	2.495	23
W_L^{\pm}	1	No	W+/W-	MW	80.4	WW	2.085	24
Z_R	1	Yes	ZR	MZR	internal	WZR	334.0	33
W_R^{\pm}	1	No	WR+/WR-	MWR	6000	WWR	190.5	34
h	0	Yes	Н	MH	125.1	WH	0.004115	25
Δ	0	Yes	DD	MDD	500	WDD	0.011	45
Η	0	Yes	HH	MHH	20000	WHH	2667.9	35
Α	0	Yes	AA	MAA	20010	WAA	2019.2	55
H^+	0	No	H+/H-	MHP	internal	WHP	2012.4	37
Δ_R^{++}	0	No	DR + +/DR-	MDRPP	500	WDRPP	25.4	9000001
Δ_R^{++} Δ_L^{++} Δ_L^{+} Δ_L^{0} χ_L^{0}	0	No	DL + +/DL-	MDLPP	internal	WDLPP	32.4	9000002
Δ_L^+	0	No	DL+/DL-	MDLP	internal	WDLP	28.8	9000003
Δ_L^0	0	Yes	DL0	MDL	500	WDL	25.5	9000004
χ_L^0	0	Yes	CHIL0	MCHIL0	internal	WCHIL0	25.5	9000005
d	$\frac{1}{2}$	No	d/d^{\sim}	MD	0.00504	0		1
и	$\frac{1}{2}$	No	u/u~	MU	0.00255	0		2
S	$\frac{1}{2}$	No	s/s~	MS	0.101	0		3
с	$\frac{1}{2}$	No	c/c~	MC	1.27	0		4
b		No	b/b^{\sim}	MB	4.7	0		5
t	$\frac{1}{2}$ $\frac{1}{2}$	No	t/t^{\sim}	MT	172	WT	1.508	6
е	$\frac{1}{2}$	No	e-/e+	ME	0.511×10^{-3}	0		11
μ	$\frac{1}{2}$	No	mu-/mu+	MMU	0.1057	0		13
τ	$\frac{1}{2}$	No	ta-/ta+	MTA	1.777	0		15
ν_1	$\frac{1}{2}$	Yes	ve	MNU1	$1. \times 10^{-12}$	0		12
ν ₂	$\frac{1}{2}$	Yes	vm	MNU2	8.9×10^{-12}	0		14
ν ₃	$\frac{1}{2}$	Yes	vt	MNU3	50.4×10^{-12}	0		16
N_1	$\frac{1}{2}$	Yes	N1	MN1	100	WN1	3.4×10^{-10}	9900012
N_2	$\frac{1}{2}$	Yes	N2	MN2	2000	WN2	5.4×10^{-4}	9900014
N ₃	$\frac{1}{2}$	Yes	N3	MN3	10000	WN3	85.9	9900016

and (101). Finally, in the NLO UFOs we removed the four-scalars interactions.

4.3 Use in Madgraph

The models were tested in Madgraph 3.5.3 both at LO and NLO order, and results for benchmark processes were obtained as reported in the next section. As a basic example one can generate the KS process at the NLO level:

```
import model mlrsm-loop
define l = l+ l-
define n = n1 n2 n3
generate p p > l n [QCD].
```

This generates $pp \rightarrow \ell N$ events, including processes mediated by W_R , by LR gauge boson mixing, and Dirac couplings, at NLO order in QCD.

A few technical notes are useful:

• In case the chosen physical inputs lead to non-perturbative couplings, Madgraph may issue a WARNING: Failed to update dependent parameter. and results will probably be invalid. We recommend checking the ranges of chose physical inputs with the LRSMEVAL function described above, or with the equivalent Python code. The python code is shipped with a jupyter notebook in which its usage is explained. Apart from providing input parameters in the form of a dictionary,

Table 4	All external	parameters ((other than	particle masses)

FR Name	FR Name Default		LHA code	Description	
tb	0.1	lrsminputs 1		Bidoublet VEVs ratio t_{β}	
alp	0.1	LRSMINPUTS 2		Bidoublet VEV v_2 phase α	
thetamix	0.01	LRSMINPUTS	3	$h - \Re \Delta_R$ mixing angle θ	
etamix	0.00001	LRSMINPUTS	4	$H - \Re \Delta_R$ mixing angle η	
phimix	0	LRSMINPUTS	5	$h - \Re \phi_{20}$ mixing angle ϕ	
r4	0	LRSMINPUTS	6	$ \rho_4 $ scalar coupling	
zetaLR	1	LRSMINPUTS	7	Ratio of Left/Right gauge couplings	
aS	0.1184	SMINPUTS	1	Strong coupling	
gw	0.65	SMINPUTS	2	Weak coupling g_L	
aEWM1	127.9	SMINPUTS	3	Inverse of electromagnetic coupling	
Gf	11.66/TeV ²	SMINPUTS	4	Fermi coupling	
MU_R	91.19 GeV	LOOP	1	NLO renormalization scale	
CKMlam	0.2245	CKMBLOCK	1		
CKMA	0.836	CKMBLOCK	2	Left CKM mixing entries	
CKMrho	0.122	CKMBLOCK	3		
CKMeta	0.355	CKMBLOCK	4		
CKMRs12	0.2245	CKMBLOCK	5		
CKMRs13	0.0036	CKMBLOCK	6	Right CKM mixing entries	
CKMRs23	0.0421	CKMBLOCK	7		
CKMRdel	1.2404	CKMBLOCK	8		
PMNSLs12	0.544	PMNSBLOCK	1		
PMNSLs23	0.756	PMNSBLOCK	2	Left PMNS mixing entries	
PMNSLs13	0.148	PMNSBLOCK	3		
PMNSLdel	3.437	PMNSBLOCK	4		
PMNSRs12	0.544	PMNSBLOCK	11		
PMNSRs23	0.756	PMNSBLOCK	12	Right PMNS mixing entries	
PMNSRs13	0.148	PMNSBLOCK	13		
PMNSRdel	3.437	PMNSBLOCK	14		
PMNSLphi1	0	PMNSBLOCK	15	Left Majorana CP phases	
PMNSLphi2	0	PMNSBLOCK	16		
PMNSRphi1	0	PMNSBLOCK	17		
PMNSRphi2	0	PMNSBLOCK	18	Right Majorana CP phases	
PMNSRphi3	0	PMNSBLOCK	19		
MNU0	0.05 eV	PMNSBLOCK	20	Lightest neutrino mass	
DMsol	$7.41 \times 10^{-5} \mathrm{eV^2}$	PMNSBLOCK	21	Δm_{sol}^2	
DMatm	$2.51 \times 10^{-3} \mathrm{eV^2}$	PMNSBLOCK	22	Δm_{atm}^2	
NUH	1	PMNSBLOCK	23	Neutrino Hierarchy (≥ 0 Normal, < 0 Inverted	

it also includes a "param_card reader" that automatically converts the parameters stored in a Madgraph param_card.dat to the internal syntax.

• For the full model files at NLO, or if neutrinos masses and Dirac couplings are not restricted to zero, update of card parameters in Madgraph hits a time constraint and the following warning is issued: WARNING: The model takes too long to load so we bypass the updating of dependent parameter In this case the user should insert an "update dependent" command just after card edition, before event generation.

• The coupling orders of all interactions are just QED or QCD, with usual hierarchy QED=2, QCD=1.

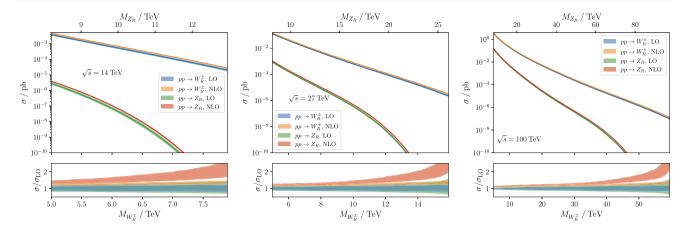


Fig. 1 Production cross-sections for W_R and Z_R at LO and NLO for $\sqrt{s} = 14, 27$ and 100 TeV. The lower part shows the *K*-factor, in which the shaded regions denote the scale- and pdf-uncertainties added in

quadrature. Note that M_{Z_R} is shown on the top x axis and explains the kinematic suppression of $pp \to Z_R$ for large M_{Z_R}

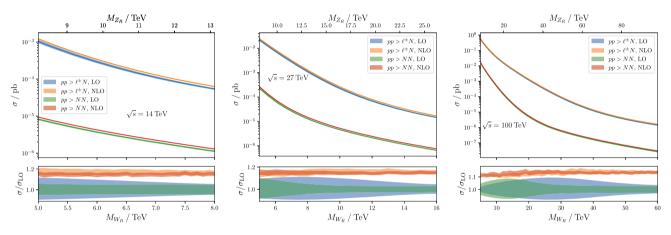


Fig. 2 Inclusive production cross-sections for $pp \rightarrow \ell^{\pm}N$ ($\ell^{\pm} = e^{\pm}, \mu^{\pm}, \tau^{\pm}, N = N_1, N_2, N_3$) and $pp \rightarrow NN$ at LO and NLO for $\sqrt{s} = 14, 27, 100$ TeV. The lower part shows the *K*-factor in which the shaded regions denote the scale- and pdf-uncertainties added in quadrature

5 Benchmarks

In this section we give an overview of a few example processes that can be consistently studied at NLO in QCD with the new model file. For all processes considered we use the following computational setup:

- We use the mlrsm-nu-loop version of the model file (5 dynamical flavours, charged lepton and light neutrino masses/yukawas have been set to zero).
- All parameters are fixed to their default values as shown in Table 4 unless otherwise indicated.
- We only compute the fixed-order cross sections without hadron shower at LO and NLO.
- We use MadGraph5_aMC_v3.5.3[130] with LHAPDF 6.5.4 [131] with default settings.
- The pdfsets are NNPDF40_nlo_as_01190 [132] (lhaid=334100).

• For the factorization and renormalization scales we use the default dynamical scale scheme, e.g. for leading order processes dynamical scale scheme #4 ($\mu_F = \mu_R = \sqrt{\hat{s}}$); pdf and scale uncertainties are added in quadrature.

In order to estimate the impact of NLO QCD corrections, we define the *K*-factor as

$$K = \frac{\sigma \pm \delta \sigma}{\sigma_{\rm LO}},\tag{147}$$

in which $\delta\sigma$ denotes the uncertainty due to pdf uncertainties (estimated from the pdf replicas) and scale variation. The factorization scale μ_F and the renormalization scale μ_R are taken dynamical as $\mu_{F,R} = S_{F,R}\sqrt{\hat{s}}$, where \hat{s} is the partonic center of mass energy and $S_{F,R}$ is a scale factor. The residual scale dependence is estimated by varying the scale factors $S_{F,R}$ independently on the interval $[\frac{1}{2}, 1, 2]$ leading to 9 different values of the cross-section. As the central value one takes $S_F = S_R = 1$, while the scale variation is quoted as

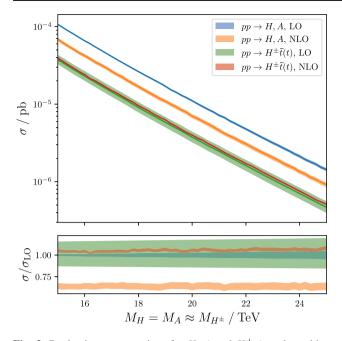


Fig. 3 Production cross-sections for H, A and H^{\pm} (together with a (anti-)top quark) for $\sqrt{s} = 100$ TeV. The lower part shows the *K*-factor in which the shaded regions denote the pdf and scale uncertainties added in quadrature

the envelope of the largest negative/positive deviation. Generally, one can expect that NLO corrections lead to a decrease of this residual scale dependence due to an order-by-order absorption of the involved scale-dependent terms.

Let us begin by studying the Drell-Yan production cross sections of the new massive gauge bosons. In Fig. 1 we show the 2 \rightarrow 1 cross-sections of $pp \rightarrow W_R^{\pm}$ and $pp \rightarrow Z_R$ for $\sqrt{s} = 14$, 27 and 100 TeV for a wide range of masses $M_{W_{P}^{\pm}}$ and $M_{Z_R} \approx 1.67 M_{W_p^{\pm}}$ (cf. Equation (25)). Note that the mass M_{Z_R} is given in secondary x-axis on top. In the lower parts of the plots we show the LO and NLO K-factors in which the shaded regions denote the uncertainty due to pdf uncertainties and scale variation. The NLO corrections lead to an increase of the cross-section of order $\mathcal{O}(10-20\%)$ (except for M_{Z_P} close to the kinematic threshold) and lead to a reduction of the scale dependence from $\mathcal{O}(10-15\%)$ to $\mathcal{O}(2-5\%)$. For larger masses M_{Z_R} , close to the kinematic threshold, the Z_R production cross-section strongly decreases while the size of the NLO corrections (and therefore the k-factors) and also scale dependence strongly increase. The production cross section strongly increases with center of mass energy by three orders of magnitude between $\sqrt{s} = 14$ TeV and $\sqrt{s} = 100$ TeV.

In Fig. 2 we show our results for the inclusive production cross sections $pp \rightarrow \ell^{\pm}N$ and $pp \rightarrow NN$, in which $\ell^{\pm} = e^{\pm}, \mu^{\pm}$ and τ^{\pm} , while $N = N_1, N_2$ and N_3 (except for $pp \rightarrow NN$ at $\sqrt{s} = 14$ TeV, since only N_1 and N_2 are kinematically accessible). Since the $\ell^{\pm}N$ (NN) production is dominantly mediated via the Drell–Yan process

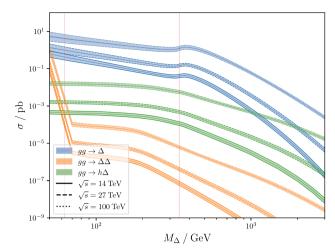


Fig. 4 Production cross-sections for $gg \rightarrow \Delta$ (blue), $gg \rightarrow \Delta\Delta$ (orange) and $gg \rightarrow h\Delta$ (green) for $\sqrt{s} = 14$, 27, 100 TeV (solid, dashed and dotted boundaries respectively). The shaded region denotes the pdf and scale uncertainty added in quadrature, while the vertical red lines denote the Higgs mass threshold $(M_{\Delta} = \frac{M_{h}}{2})$ and the top mass threshold $(M_{\Delta} = 2M_{t})$. All cross sections are calculated for $\theta = 0.1$

 $pp \rightarrow W_R^{\pm}(Z_R) \rightarrow \ell^{\pm}N(NN)$, we observe a very similar impact of the NLO corrections as for the $pp \rightarrow W_R^{\pm}$, Z_R production cross-sections for most values of M_{W_R} and M_{Z_R} . For larger M_{Z_R} , such that the on-shell production of Z_R as a resonance is suppressed (see also Fig. 1), the cross-section is dominated by non-resonant NN-production, leading to a different behavior of the cross-sections. With our choice of the heavy neutrino masses (cf. Table 4), both cross sections are dominated by having N_1 in the final state.

The Drell–Yan production cross-sections for heavy (pseudo-) scalars A and H as well as the heavy charged scalar H^{\pm} (with an additional (anti-) top quark in the final state) proceeds mostly via initial state b-quarks and a gluon, while gluon initiated processes give a 2–3 order of magnitude smaller contribution. In Fig. 3 we show our results for the $pp \rightarrow H^{\pm}\bar{t}(t)$ production cross-sections for $\sqrt{s} = 100$ TeV. The masses M_H and M_A are set equal, while the mass $M_{H^{\pm}}$ is degenerate with M_A up to order ϵ^2 corrections.

Firstly, we notice that for $pp \rightarrow A$, H the NLO corrections lead to a significant *decrease* of the cross section by 30–40%, due to reduced high scale top Yukawa, while the scale dependence slightly increases. The NLO corrections $pp \rightarrow H^{\pm}\bar{t}(t)$ lead to a mild increase of the cross section and a significant reduction of the scale variation from $\mathcal{O}(15-20\%)$ to $\mathcal{O}(1-2\%)$.

Lastly, we consider several production mechanisms of the neutral scalar Δ . Since Δ mostly comes from the real part of the neutral component of the right-handed triplet Δ_R , it couples to quarks only through mixing with with the SM-like scalar *h*. Therefore, its largest quark coupling is to the top quark, making different type of gluon fusion pro-

Table 5Example cross sections for the set of input parameters given in Table 4. The (super)sub-scripts denote the (upper)lower envelope due topdf and scale uncertainties. We also quote the central K-factor

Process	$\sigma_{ m LO}$ [fb]	$\sigma_{ m NLO}$ [fb]	K	\sqrt{s} [TeV]
	$0.636^{+16\%}_{-13\%}$	$0.806^{+4.4\%}_{-5.4\%}$	1.27	14
$pp \rightarrow W_R^{\pm}$	$38.21^{+11\%}_{-9.4\%}$	$46.0^{+2.6\%}_{-3.2\%}$	1.20	27
	$1572^{+3.2\%}_{-3.0\%}$	$1791^{+1.2\%}_{-1.2\%}$	1.14	100
	$7.0^{+26\%}_{-20\%}\times10^{-5}$	$10.6^{+10\%}_{-10\%}\times10^{-5}$	1.52	14
$pp \rightarrow Z_R$	$0.16^{+14\%}_{-12\%}$	$0.20^{+3.5\%}_{-4.4\%}$	1.25	27
	$59.9^{+5.7\%}_{-5.2\%}$	$68.6^{+1.5\%}_{-1.6\%}$	1.15	100
	$0.29^{+10\%}_{-8\%}$	$0.34^{+2.3\%}_{-3.0\%}$	1.17	14
$pp \to \ell^{\pm} N$	$2.79^{+11\%}_{-9\%}$	$3.22^{+1.8\%}_{-2.7\%}$	1.15	27
	$271.0^{+3.0\%}_{-2.9\%}$	$303.3^{+1.2\%}_{-1.1\%}$	1.12	100
	$3.72^{+5.0\%}_{-4.5\%} \times 10^{-3}$	$4.24^{+1.7\%}_{-1.8\%} \times 10^{-3}$	1.14	14
$pp \rightarrow NN$	$6.7^{+8\%}_{-7\%} \times 10^{-2}$	$7.75^{+1.8\%}_{-2.4\%} imes 10^{-2}$	1.15	27
	6.8 ^{+5.6%} -5.1%	$7.59^{+1.1\%}_{-1.3\%}$	1.11	100
$pp \to H$	$8.1^{+21\%}_{-16\%} \times 10^{-12}$	$7.2^{+2.2\%}_{-5.2\%} \times 10^{-12}$	0.89	27
	$1.11^{+2.9\%}_{-3.2\%} \times 10^{-2}$	$7.0^{+5.7\%}_{-6.0\%} \times 10^{-3}$	0.63	100
$pp \to A$	$8.0^{+21\%}_{-16\%} \times 10^{-12}$	$7.1^{+2.2\%}_{-5.2\%} \times 10^{-12}$	0.89	27
	$1.10^{+2.9\%}_{-3.2\%} \times 10^{-2}$	$7.0^{+5.7\%}_{-6.0\%}10^{-3}$	0.63	100
$pp \to H^{\pm} \bar{t}(t)$	$8.0^{+36\%}_{-25\%} \times 10^{-13}$	$1.0^{+7\%}_{-10\%} \times 10^{-12}$	1.35	27
••	$3.8^{+18\%}_{-14\%} \times 10^{-3}$	$4.0^{+1.2\%}_{-2.3\%} \times 10^{-3}$	1.06	100
	$0.19^{+26\%}_{-20\%}$		_	14
$gg \to \Delta$	$0.84^{+21\%}_{-17\%}$	_	_	27
	$8.6^{+25\%}_{-20\%}$	_	_	100
	$3.7^{+31\%}_{-22\%} \times 10^{-9}$	_	_	14
$gg \to \Delta \Delta$	$2.7^{+26\%}_{-19\%} imes 10^{-8}$	_	_	27
	$5.1^{+17\%}_{-14\%} \times 10^{-7}$	_	_	100
	$3.2^{+29\%}_{-21\%} \times 10^{-4}$	_	_	14
$gg \to \Delta h$	$1.7^{+23\%}_{-18\%} \times 10^{-3}$	_	_	27
	$2.4^{+20\%}_{-16\%} \times 10^{-2}$	_	_	100

cesses the dominant production mechanisms. Since gluon fusion is a loop-induced process, we can only consider the leading order cross-sections within the scope of Madgraph and this model file. In particular, we consider single- and pair production via gluon fusion $gg \rightarrow \Delta, \Delta\Delta$ as well as "Higgs- Δ -strahlung" $gg \rightarrow h\Delta$ which proceeds both via $gg \rightarrow h \rightarrow h\Delta$, as well as $gg \rightarrow \Delta \rightarrow h\Delta$ in addition to box-topologies. In Fig.4 we show our results for these production mechanisms for $m_{\Delta} \in (50, 3000)$ GeV for $\sqrt{s} = 14,27$ and 100 TeV. In addition, we set here $\theta = 0.1$ to attain sizeable production cross sections. Firstly, we notice that for $M_{\Delta} \simeq 2m_t$ (denoted by the right vertical red line) the $gg \rightarrow \Delta$ cross section increases due to threshold effects in the corresponding "top-triangle", before it drops off for larger masses. Next, for $M_{\Delta} \lesssim M_h/2$ (denoted by the left vertical red line), the $gg \rightarrow \Delta \Delta$ production cross section becomes large due to an on-shell intermediate Higgs in $gg \rightarrow h \rightarrow \Delta \Delta$. Above threshold, the cross-section strongly decreases and becomes somewhat subdominant. Lastly, the process $gg \rightarrow h\Delta$ has a sizeable cross section for a wide range of masses $M_{\Delta} \in (M_h/2, \sim 500)$ GeV, competing with $gg \rightarrow \Delta$ for larger masses.

In addition to the aforementioned "benchmark processes", we also checked several SM processes with the LRSM model file, such as $pp \rightarrow t\bar{t}$ and $gg \rightarrow h \rightarrow t\bar{t}$ production cross-sections, as well as $pp \rightarrow VV$ ($V = W^{\pm}, Z$), finding perfect agreement with the built-in Madgraph models. Finally, in Table 5 we summarize our results for the choice of parameters given in Table 4 as a representative benchmark point. This phenomenological study is by no means complete and rather serves as a proof of concept. We would like to emphasize that with our model file fully consistent collider stud-

ies (cross-section calculation and event generation) are now possible with NLO QCD corrections as well as the study of loop-induced processes at LO. The NLO corrections generally lead to a significant decrease of the scale dependence and therefore to a significant reduction of signal modelling uncertainties for collider studies.

6 Summary and outlook

The main purpose of this work was to provide an implementation of the LRSM, which is as complete as possible and can be used for a variety of collider and other phenomenological studies. The collider frontier in the near future will mostly be focused on the HL-LHC, but in the less near future we may expect to see lepton colliders FCC-ee, ILC, CLIC or muon colliders, and finally high energetic scatterings of protons in FCC-eh or FCC-hh at few tens of TeV (30 or 100 TeV). With the present model file, the NLO QCD corrections can be reliably calculated for any process within the LRSM, properly including the mixings of scalars, gauge bosons and fermions in different gauges, too. Not all the parts of the model are necessarily needed for simulations, even very precise ones, therefore we provide a number of restricted versions, which may be faster and easier to use.

Perhaps the most significant analytical developments in this work are the explicit calculation of heavy-light neutrino mixing and the solution of the scalar spectrum. For neutrinos, the Majorana-Dirac connection was found some time ago [67], but with the help of the Cayley–Hamilton theorem, we managed to come up with a closed form solution for a general complex matrix in terms of its invariants. This relation may become very useful to clarify the flavor relations in the leptonic sector, e.g. how the Dirac induced transitions (in eEDMs, $0v2\beta$, collider of $N_1 \rightarrow v\gamma$) are related to the flavor structure of V_R and the m_N mass spectrum. The explicit result allows for a direct insight by taking certain expansions (small angles or masses) or correction to limits, like type II (sub)dominance.

In the scalar sector, the mass spectrum was derived for the first time without extra assumptions of small mixing angles, thus performing an exact diagonalization. The only phenomenologically required assumption was that the Higgs field *h* has a limited admixture with heavy scalars. Remarkably, we were able to derive exact expressions for the model couplings in terms of scalar masses and mixings, which can cover both the standard regime of heavy scalars as well as the regime of light Δ . The latter case is relevant for the Majorana–Higgs program, but the provided solution covers the regimes of quasi-degeneracy, too. Criteria for perturbativity of the chosen inputs were explicitly identified, and provided as an implemented routine. The model capabilities were demonstrated with a number of benchmarks concerning the single production of heavy resonances, $pp \rightarrow W_R$, Z_R , various scalars $pp \rightarrow \Delta$, H, A, H^{\pm} and production of heavy neutrinos $pp \rightarrow \ell N$, $pp \rightarrow NN$, calculated at the NLO level in QCD.

Another foreseeable improvement might be the inclusion of electroweak corrections. These could be used to automatize the computation of electroweak precision observables and calculate one-loop processes, like radiative production and decays at one loop. Such an analysis may not be impossible but it would require going beyond the current state of the art, given the complexity of the model. We shall leave it for a future study. The payoff might be having precise and automatized control over rare processes that might be possible to explore at future colliders.

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Data Availability Statement My manuscript has no associated data. [Authors' comment: Data sharing not applicable to this article as no datasets were generated or analysed during the current study.]

Code Availability Statement My manuscript has code/software included as electronic supplementary material. [Authors' comment: All code/software generated or analysed during this study is included in this published article [and its supplementary information files].]

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Funded by SCOAP³.

Appendix A: Square root(s) of a 3 × 3 matrix

We calculate the closed form analytic expression for the 8 square roots of a general complex 3×3 matrix using the Cayley–Hamilton theorem. For some of the previous work, see [133–135].

We start by writing the invariants of A and \sqrt{A}

$$T = \operatorname{tr} A, \qquad \qquad T_{1/2} = \operatorname{tr} \sqrt{A}, \qquad (A1)$$

$$T_2 = \text{tr } A.A$$
, $T_{2,1/2} = \text{tr } \sqrt{A}.\sqrt{A} = T$, (A2)

$$\Delta = \det A$$
, $\Delta_{1/2} = \det \sqrt{A} = \sqrt{\Delta}$, (A3)

where the trace of the square of \sqrt{A} and its determinant are trivially related to the invariants of A. Cayley–Hamilton theorem states that matrices themselves are solutions to their characteristic polynomial equations, which can be written in terms of invariants

$$A^{3} - TA^{2} + \tilde{T}A - \Delta \mathbb{1} = 0, \quad \tilde{T} = \frac{1}{2} \left(T^{2} - T_{2} \right).$$
 (A4)

The same holds for the square root that obeys a polynomial with its own invariants (two of which are directly related to the invariants of A)

$$\left(\sqrt{A}\right)^3 - T_{1/2}\left(\sqrt{A}\right)^2 + \tilde{T}_{1/2}\sqrt{A} - \sqrt{\Delta}\,\mathbb{1} = 0\,,\qquad(A5)$$

$$\tilde{T}_{1/2} = \frac{1}{2} \left(T_{1/2}^2 - T \right), \tag{A6}$$

$$A\sqrt{A} - T_{1/2}A + \tilde{T}_{1/2}\sqrt{A} - \sqrt{\Delta} \mathbb{1} = 0.$$
 (A7)

We can get to a linear equation for \sqrt{A} by multiplying (A7) with \sqrt{A}

$$A^{2} - T_{1/2}A\sqrt{A} + \tilde{T}_{1/2}A - \sqrt{\Delta}\sqrt{A} = 0, \qquad (A8)$$

$$A^{2} - T_{1/2} \left(T_{1/2}A - \tilde{T}_{1/2}\sqrt{A} + \sqrt{\Delta} \mathbb{1} \right) + \tilde{T}_{1/2}A - \sqrt{\Delta}\sqrt{A} = 0, \quad (A9)$$

which finally gives

$$\sqrt{A} = \pm \frac{A^2 + \left(\tilde{T}_{1/2} - T_{1/2}^2\right)A - \sqrt{\Delta} T_{1/2}\mathbb{1}}{\sqrt{\Delta} - T_{1/2}\tilde{T}_{1/2}}.$$
 (A10)

This is already a striking result, where we know the offdiagonal flavor structure of the root directly from a power expansion of A, terminating at the second order (as expected for 3×3 matrices from the Cayley–Hamilton theorem).

All the non-linearity is hidden in the equation for the remaining invariant $T_{1/2}$, which we get by taking the trace of (A9) and using (A6), such that

$$T_{1/2}^4 - 2T T_{1/2}^2 - 8\sqrt{\Delta} T_{1/2} + 2T_2 - T^2 = 0.$$
 (A11)

This is a depressed quartic with four possible solutions

$$T_{1/2} = \frac{\pm \eta_s + s \,\chi}{2\sqrt{6\xi}}, \qquad s = \pm, \qquad (A12)$$

which in turn give four corresponding $\tilde{T}_{1/2}$ via (A6).

The shorthands ξ , χ , η_{\pm} , are only functions of the invariants of *A* and are found from these expressions:

$$\xi^{3} = -32 \left(5T^{3} - 9TT_{2} - 54\Delta \right) + 96\sqrt{3 \left(T^{2} - 2T_{2}\right) \left(T^{2} - T_{2}\right)^{2} + 12T \left(9T_{2} - 5T^{2}\right) \Delta} + 324\Delta^{2}, \qquad (A13)$$

$$\chi^{2} = -16\sqrt[3]{2}T^{2} + 48\sqrt[3]{2}T_{2} + 8T\xi + \sqrt[3]{4}\xi^{2}, \qquad (A14)$$

$$\eta_{\pm}^{2} = 16 \sqrt[3]{2} \left(T^{2} - 3T_{2} \right) + \xi \left(16T - \sqrt[3]{4} \xi \pm 96\sqrt{6} \chi^{-1} \sqrt{\Delta \xi} \right),$$
(A15)

where one assumes positive roots only. Together with the overall sign in (A10), we get the total of 8 matrix roots.

With an explicit expression for \sqrt{A} , it is also easy to get the inverse root, by multiplying (A7) with $A^{-1/2}$,

$$A - T_{1/2}\sqrt{A} + \tilde{T}_{1/2} \mathbb{1} - \sqrt{\Delta} A^{-1/2} = 0, \qquad (A16)$$

so that the inverse root is found as

$$\left(\sqrt{A}\right)^{-1} = \frac{1}{\sqrt{\Delta}} \left(A - T_{1/2}\sqrt{A} + \tilde{T}_{1/2} \mathbb{1} \right).$$
 (A17)

Again, there are eight possible branches and also the form of this solution is a matrix power expansion up to A^2 .

References

- S. Weinberg, Phys. Rev. Lett. 19, 1264 (1967). https://doi.org/10. 1103/PhysRevLett.19.1264
- G. Aad et al. (ATLAS), Nature 607, 52 (2022). https://doi.org/10. 1038/s41586-022-04893-w. arXiv:2207.00092 [hep-ex]. [Erratum: Nature 612, E24 (2022)]
- A. Tumasyan et al. (CMS), Nature 607, 60 (2022). https://doi.org/ 10.1038/s41586-022-04892-x. arXiv:2207.00043 [hep-ex]
- J.C. Pati, A. Salam, Phys. Rev. D 10, 275 (1974). https://doi. org/10.1103/PhysRevD.10.275. [Erratum: Phys. Rev. D 11, 703 (1975)]
- R.N. Mohapatra, J.C. Pati, Phys. Rev. D 11, 2558 (1975). https:// doi.org/10.1103/PhysRevD.11.2558
- G. Senjanović, R.N. Mohapatra, Phys. Rev. D 12, 1502 (1975). https://doi.org/10.1103/PhysRevD.12.1502
- G. Senjanović, Nucl. Phys. B 153, 334 (1979). https://doi.org/10. 1016/0550-3213(79)90604-7
- R.N. Mohapatra, G. Senjanović, Phys. Rev. Lett. 44, 912 (1980).
 [231 (1979)]. https://doi.org/10.1103/PhysRevLett.44.912
- R.N. Mohapatra, G. Senjanović, Phys. Rev. Lett. 44, 231 (1979). https://doi.org/10.1103/PhysRevLett.44.912
- G. Senjanović, P. Senjanović, Phys. Rev. D 21, 3253 (1980). https://doi.org/10.1103/PhysRevD.21.3253
- P. Minkowski, Phys. Lett. 67B, 421 (1977). https://doi.org/10. 1016/0370-2693(77)90435-X
- R.N. Mohapatra, G. Senjanović, Phys. Rev. D 23, 165 (1981). https://doi.org/10.1103/PhysRevD.23.165

- M. Gell-Mann, P. Ramond, R. Slansky, Conf. Proc. C 790927, 315 (1979). arXiv:1306.4669 [hep-th]
- S.L. Glashow, NATO Sci. Ser. B 61, 687 (1980). https://doi.org/ 10.1007/978-1-4684-7197-7_15
- G. Beall, M. Bander, A. Soni, 11th International Symposium on Lepton and Photon Interactions at High Energies Ithaca, New York, August 4–9, 1983. Phys. Rev. Lett. 48, 848 (1982). https:// doi.org/10.1103/PhysRevLett.48.848
- R.N. Mohapatra, G. Senjanovic, M.D. Tran, Phys. Rev. D 28, 546 (1983). https://doi.org/10.1103/PhysRevD.28.546
- G. Ecker, W. Grimus, Nucl. Phys. B 258, 328 (1985). https://doi. org/10.1016/0550-3213(85)90616-9
- Y. Zhang, H. An, X. Ji, R.N. Mohapatra, Nucl. Phys. B 802, 247 (2008). https://doi.org/10.1016/j.nuclphysb.2008.05. 019. arXiv:0712.4218 [hep-ph]
- Y. Zhang, H. An, X. Ji, R.N. Mohapatra, Phys. Rev. D 76, 091301 (2007). https://doi.org/10.1103/PhysRevD.76.091301. arXiv:0704.1662 [hep-ph]
- A. Maiezza, M. Nemevšek, F. Nesti, G. Senjanović, Phys. Rev. D 82, 055022 (2010). https://doi.org/10.1103/PhysRevD.82. 055022. arXiv:1005.5160 [hep-ph]
- 21. S. Bertolini, A. Maiezza, F. Nesti, Phys. Rev. D 88, 034014 (2013). https://doi.org/10.1103/PhysRevD.88.034014. arXiv:1305.5739 [hep-ph]
- A. Maiezza, M. Nemevšek, Phys. Rev. D 90, 095002 (2014). https://doi.org/10.1103/PhysRevD.90.095002. arXiv:1407.3678 [hep-ph]
- S. Bertolini, A. Maiezza, F. Nesti, Phys. Rev. D 101, 035036 (2020). https://doi.org/10.1103/PhysRevD.101.035036. arXiv:1911.09472 [hep-ph]
- S. Bertolini, L. Di Luzio, F. Nesti, Phys. Rev. Lett. 126, 081801 (2021). https://doi.org/10.1103/PhysRevLett.126. 081801. arXiv:2006.12508 [hep-ph]
- V. Bernard, S. Descotes-Genon, L. Vale Silva, JHEP 2020(09), 088 (2020). https://doi.org/10.1007/JHEP09(2020)088. arXiv:2001.00886 [hep-ph]
- W. Dekens, L. Andreoli, J. de Vries, E. Mereghetti, F. Oosterhof, JHEP 2021(11), 127 (2021). https://doi.org/10.1007/ JHEP11(2021)127. arXiv:2107.10852 [hep-ph]
- A. Atre, T. Han, S. Pascoli, B. Zhang, JHEP 2009(05), 030 (2009). https://doi.org/10.1088/1126-6708/2009/05/030. arXiv:0901.3589 [hep-ph]
- P.S. Bhupal Dev, R. Franceschini, R.N. Mohapatra, Phys. Rev. D 86, 093010 (2012). https://doi.org/10.1103/PhysRevD.86. 093010. arXiv:1207.2756 [hep-ph]
- Y. Cai, T. Han, T. Li, R. Ruiz, Front. Phys. 6, 40 (2018). https:// doi.org/10.3389/fphy.2018.00040. arXiv:1711.02180 [hep-ph]
- A.M. Abdullahi et al., J. Phys. G 50, 020501 (2023). https://doi. org/10.1088/1361-6471/ac98f9. arXiv:2203.08039 [hep-ph]
- W.-Y. Keung, G. Senjanović, Phys. Rev. Lett. 50, 1427 (1983). https://doi.org/10.1103/PhysRevLett.50.1427
- M. Nemevšek, F. Nesti, G. Senjanović, Y. Zhang, Phys. Rev. D 83, 115014 (2011). https://doi.org/10.1103/PhysRevD.83. 115014. arXiv:1103.1627 [hep-ph]
- M. Nemevšek, F. Nesti, G. Popara, Phys. Rev. D 97, 115018 (2018). https://doi.org/10.1103/PhysRevD.97.115018. arXiv:1801.05813 [hep-ph]
- 34. M. Mitra, R. Ruiz, D.J. Scott, M. Spannowsky, Phys. Rev. D 94, 095016 (2016). https://doi.org/10.1103/PhysRevD.94. 095016. arXiv:1607.03504 [hep-ph]
- 35. A. Dey, R. Rahaman, S.K. Rai, Eur. Phys. J. C 84, 132 (2024). https://doi.org/10.1140/epjc/s10052-024-12493-3. arXiv:2207.06857 [hep-ph]
- J. Alimena et al., J. Phys. G 47, 090501 (2020). https://doi.org/ 10.1088/1361-6471/ab4574. arXiv:1903.04497 [hep-ex]

- P. Langacker, S.U. Sankar, Phys. Rev. D 40, 1569 (1989). https:// doi.org/10.1103/PhysRevD.40.1569
- M. Frank, O. Özdal, P. Poulose, Phys. Rev. D 99, 035001 (2019). https://doi.org/10.1103/PhysRevD.99.035001. arXiv:1812.05681 [hep-ph]
- 39. S.F. Solera, A. Pich, L. Vale Silva, JHEP 2024(02), 027 (2024). https://doi.org/10.1007/JHEP02(2024)027. arXiv:2309.06094 [hep-ph]
- T.G. Rizzo, JHEP 2007(05), 037 (2007). https://doi.org/10.1088/ 1126-6708/2007/05/037. arXiv:0704.0235 [hep-ph]
- T. Han, I. Lewis, R. Ruiz, Z.-g. Si, Phys. Rev. D 87, 035011 (2013). https://doi.org/10.1103/PhysRevD.87.035011. arXiv:1211.6447 [hep-ph]. [Erratum: Phys. Rev. D 87, 039906 (2013)]
- P.S.B. Dev, D. Kim, R.N. Mohapatra, JHEP 2016(01), 118 (2016). https://doi.org/10.1007/JHEP01(2016)118. arXiv:1510.04328 [hep-ph]
- J. Gluza, T. Jelinski, R. Szafron, Phys. Rev. D 93, 113017 (2016). https://doi.org/10.1103/PhysRevD.93.113017. arXiv:1604.01388 [hep-ph]
- 44. M. Frank, B. Fuks, A. Jueid, S. Moretti, O. Ozdal, (2023). arXiv:2312.08521 [hep-ph]
- 45. S. Mandal, M. Mitra, N. Sinha, Phys. Rev. D 96, 035023 (2017). https://doi.org/10.1103/PhysRevD.96.035023. arXiv:1705.01932 [hep-ph]
- R.M. Godbole, S.P. Maharathy, S. Mandal, M. Mitra, N. Sinha, Phys. Rev. D 104, 095009 (2021). https://doi.org/10.1103/ PhysRevD.104.095009. arXiv:2008.05467 [hep-ph]
- G.F.S. Alves, C.S. Fong, L.P.S. Leal, R. Zukanovich Funchal, Phys. Rev. Lett. 133, 161802 (2024). https://doi.org/10.1103/ PhysRevLett.133.161802. arXiv:2307.04862 [hep-ph]
- R. Ruiz, Eur. Phys. J. C 77, 375 (2017). https://doi.org/10.1140/ epjc/s10052-017-4950-2. arXiv:1703.04669 [hep-ph]
- M. Nemevšek, F. Nesti, Phys. Rev. D 108, 015030 (2023). https:// doi.org/10.1103/PhysRevD.108.015030. arXiv:2306.12104 [hep-ph]
- 50. T.G. Rizzo, Phys. Rev. D **89**, 095022 (2014). https://doi.org/10. 1103/PhysRevD.89.095022. arXiv:1403.5465 [hep-ph]
- 51. J.N. Ng, A. de la Puente, B.W.-P. Pan, JHEP **2015**(12), 172 (2015). https://doi.org/10.1007/JHEP12(2015)172. arXiv:1505.01934 [hep-ph]
- T. Golling et al., (2016). https://doi.org/10.23731/ CYRM-2017-003.441. arXiv:1606.00947 [hep-ph]
- K.A. Urquía-Calderón, Phys. Rev. D 109, 055002 (2024). https:// doi.org/10.1103/PhysRevD.109.055002. arXiv:2310.17406 [hep-ph]
- A. Tumasyan et al. (CMS), JHEP 2022(04), 047 (2022). https:// doi.org/10.1007/JHEP04(2022)047. arXiv:2112.03949 [hep-ex]
- 55. G. Aad et al. (ATLAS), (2023). arXiv:2304.09553 [hep-ex]
- G. Aad et al. (ATLAS), Phys. Rev. D 100, 052013 (2019). https:// doi.org/10.1103/PhysRevD.100.052013. arXiv:1906.05609 [hep-ex]
- A. Tumasyan et al. (CMS), JHEP 2022 (07), 067 (2022). https:// doi.org/10.1007/JHEP07(2022)067. arXiv:2202.06075 [hep-ex]
- M. Aaboud et al. (ATLAS), Phys. Rev. Lett. **120**, 161802 (2018). https://doi.org/10.1103/PhysRevLett.120.161802. arXiv:1801.06992 [hep-ex]
- 59. A. Collaboration (ATLAS), ATLAS-CONF (2021)
- A. Tumasyan et al. (CMS), JHEP 2023 (09), 051 (2023). https:// doi.org/10.1007/JHEP09(2023)051. arXiv:2212.12604 [hep-ex]
- 61. G. Aad et al. (ATLAS), JHEP **2020**(03), 145 (2020). https://doi. org/10.1007/JHEP03(2020)145. arXiv:1910.08447 [hep-ex]
- A.M. Sirunyan et al. (CMS), JHEP 2020(05), 033 (2020). https:// doi.org/10.1007/JHEP05(2020)033. arXiv:1911.03947 [hep-ex]
- 63. A. Hayrapetyan et al. (CMS), (2023). arXiv:2310.19893 [hep-ex]
- G. Aad et al. (ATLAS), JHEP 2023(12), 073 (2023). https://doi. org/10.1007/JHEP12(2023)073. arXiv:2308.08521 [hep-ex]

- 65. G. Aad et al. (ATLAS), JHEP **2023**(06), 016 (2023). https://doi. org/10.1007/JHEP06(2023)016. arXiv:2207.00230 [hep-ex]
- 66. C. Collaboration (CMS), (2022). arXiv:2210.00043 [hep-ex]
- M. Nemevšek, G. Senjanović, V. Tello, Phys. Rev. Lett. 110, 151802 (2013). https://doi.org/10.1103/PhysRevLett.110. 151802. arXiv:1211.2837 [hep-ph]
- C.-Y. Chen, P.S.B. Dev, R.N. Mohapatra, Phys. Rev. D 88, 033014 (2013). https://doi.org/10.1103/PhysRevD.88.033014. arXiv:1306.2342 [hep-ph]
- 69. C. Arbelaéz, C. Dib, I. Schmidt, J.C. Vasquez, Phys. Rev. D 97, 055011 (2018). https://doi.org/10.1103/PhysRevD.97. 055011. arXiv:1712.08704 [hep-ph]
- 70. G. Li, M.J. Ramsey-Musolf, J.C. Vasquez, Phys. Rev. D 105, 115021 (2022). https://doi.org/10.1103/PhysRevD.105.115021. arXiv:2202.01789 [hep-ph]
- V. Tello, M. Nemevšek, F. Nesti, G. Senjanović, F. Vissani, Phys. Rev. Lett. **106**, 151801 (2011). https://doi.org/10.1103/ PhysRevLett.106.151801. arXiv:1011.3522 [hep-ph]
- M. Nemevšek, F. Nesti, G. Senjanović, V. Tello, (2011). arXiv:1112.3061 [hep-ph]
- J. Barry, W. Rodejohann, J. High Energy Phys. 2013, 153 (2013). https://doi.org/10.1007/JHEP09(2013)153
- 74. J. de Vries, G. Li, M.J. Ramsey-Musolf, J.C. Vasquez, JHEP 2022(11), 056 (2022). https://doi.org/10.1007/ JHEP11(2022)056. arXiv:2209.03031 [hep-ph]
- M. Nemevšek, G. Senjanović, Y. Zhang, JCAP 2012(07), 006 (2012). https://doi.org/10.1088/1475-7516/2012/07/006. arXiv:1205.0844 [hep-ph]
- 76. M. Nemevšek, Y. Zhang, Phys. Rev. Lett. 130, 121002 (2023). https://doi.org/10.1103/PhysRevLett.130.121002. arXiv:2206.11293 [hep-ph]
- 77. M. Nemevšek, Y. Zhang, (2023). arXiv:2312.00129 [hep-ph]
- 78. G. Senjanović, V. Tello, Phys. Rev. D 100, 115031 (2019). https:// doi.org/10.1103/PhysRevD.100.115031. arXiv:1812.03790 [hep-ph]
- 79. G. Senjanović, V. Tello, (2020). arXiv:2004.04036 [hep-ph]
- J. Kiers, K. Kiers, A. Szynkman, T. Tarutina, Phys. Rev. D 107, 075001 (2023). https://doi.org/10.1103/PhysRevD.107.075001. arXiv:2212.14837 [hep-ph]
- F.I. Olness, M.E. Ebel, Phys. Rev. D 32, 1769 (1985). https://doi. org/10.1103/PhysRevD.32.1769
- J.F. Gunion, J. Grifols, A. Mendez, B. Kayser, F.I. Olness, Phys. Rev. D 40, 1546 (1989). https://doi.org/10.1103/PhysRevD.40. 1546
- N.G. Deshpande, J.F. Gunion, B. Kayser, F.I. Olness, Phys. Rev. D 44, 837 (1991). https://doi.org/10.1103/PhysRevD.44.837
- K. Kiers, M. Assis, A.A. Petrov, Phys. Rev. D 71, 115015 (2005). https://doi.org/10.1103/PhysRevD.71.115015. arXiv:hep-ph/0503115
- O. Khasanov, G. Perez, Phys. Rev. D 65, 053007 (2002). https:// doi.org/10.1103/PhysRevD.65.053007. arXiv:hep-ph/0108176
- W. Dekens, D. Boer, Nucl. Phys. B 889, 727 (2014). https://doi. org/10.1016/j.nuclphysb.2014.10.025. arXiv:1409.4052 [hep-ph]
- G. Bambhaniya, J. Chakrabortty, J. Gluza, T. Jelinski, R. Szafron, Phys. Rev. D 92, 015016 (2015). https://doi.org/10.1103/ PhysRevD.92.015016. arXiv:1504.03999 [hep-ph]
- A. Maiezza, G. Senjanović, J.C. Vasquez, Phys. Rev. D 95, 095004 (2017). https://doi.org/10.1103/PhysRevD.95.095004. arXiv:1612.09146 [hep-ph]
- P.S.B. Dev, R.N. Mohapatra, Y. Zhang, JHEP 2016(05), 174 (2016). https://doi.org/10.1007/JHEP05(2016)174. arXiv:1602.05947 [hep-ph]
- 90. A. Maiezza, M. Nemevšek, F. Nesti, Phys. Rev. D 94, 035008 (2016). https://doi.org/10.1103/PhysRevD.94.035008. arXiv:1603.00360 [hep-ph]

- 91. G. Chauhan, P.S.B. Dev, R.N. Mohapatra, Y. Zhang, JHEP **2019**(01), 208 (2019). https://doi.org/10.1007/ JHEP01(2019)208. arXiv:1811.08789 [hep-ph]
- R.N. Mohapatra, G. Yan, Y. Zhang, Nucl. Phys. B 948, 114764 (2019). https://doi.org/10.1016/j.nuclphysb.2019.114764. arXiv:1902.08601 [hep-ph]
- M. Nemevšek, F. Nesti, J.C. Vasquez, JHEP 2017(04), 114 (2017). https://doi.org/10.1007/JHEP04(2017)114. arXiv:1612.06840 [hep-ph]
- 94. P.S. Bhupal Dev, R.N. Mohapatra, W. Rodejohann, X.-J. Xu, JHEP **2019**(02), 154 (2019). https://doi.org/10.1007/ JHEP02(2019)154. arXiv:1811.06869 [hep-ph]
- G. Chauhan, JHEP 2019(12), 137 (2019). https://doi.org/10.1007/ JHEP12(2019)137. arXiv:1907.07153 [hep-ph]
- 96. V. Brdar, L. Graf, A.J. Helmboldt, X.-J. Xu, JCAP **2019**(12), 027 (2019). https://doi.org/10.1088/1475-7516/2019/12/027. arXiv:1909.02018 [hep-ph]
- 97. M. Li, Q.-S. Yan, Y. Zhang, Z. Zhao, JHEP **2021**(03), 267 (2021). https://doi.org/10.1007/JHEP03(2021)267. arXiv:2012.13686 [hep-ph]
- A. Maiezza, M. Nemevšek, F. Nesti, Phys. Rev. Lett. 115, 081802 (2015). https://doi.org/10.1103/PhysRevLett.115. 081802. arXiv:1503.06834 [hep-ph]
- A. Melfo, M. Nemevsek, F. Nesti, G. Senjanovic, Y. Zhang, Phys. Rev. D 85, 055018 (2012). https://doi.org/10.1103/PhysRevD.85. 055018. arXiv:1108.4416 [hep-ph]
- 100. J. Maalampi, N. Romanenko, Phys. Lett. B 532, 202 (2002). https://doi.org/10.1016/S0370-2693(02)01549-6. arXiv:hep-ph/0201196
- 101. A. Roitgrund, G. Eilam, JHEP **2021**(01), 031 (2021). https:// doi.org/10.1007/JHEP01(2021)031. arXiv:1704.07772 [hep-ph]. [Erratum: JHEP **03**, 029 (2021)]
- 102. A. Alloul, M. Frank, B. Fuks, M. Rausch de Traubenberg, Phys. Rev. D 88, 075004 (2013). https://doi.org/10.1103/PhysRevD.88. 075004. arXiv:1307.1711 [hep-ph]
- 103. E. Akhmedov, P.S.B. Dev, S. Jana, R.N. Mohapatra, (2024). arXiv:2401.15145 [hep-ph]
- 104. S. Alekhin et al., Rep. Prog. Phys. **79**, 124201 (2016). https://doi. org/10.1088/0034-4885/79/12/124201. arXiv:1504.04855 [hepph]
- 105. F. Kling, S. Trojanowski, Phys. Rev. D 97, 095016 (2018). https:// doi.org/10.1103/PhysRevD.97.095016. arXiv:1801.08947 [hepph]
- 106. N.D. Christensen, C. Duhr, Comput. Phys. Commun. 180, 1614 (2009). https://doi.org/10.1016/j.cpc.2009.02.018. arXiv:0806.4194 [hep-ph]
- 107. A. Alloul, N.D. Christensen, C. Degrande, C. Duhr, B. Fuks, Comput. Phys. Commun. 185, 2250 (2014). https://doi.org/10.1016/j. cpc.2014.04.012. arXiv:1310.1921 [hep-ph]
- C. Degrande, C. Duhr, B. Fuks, D. Grellscheid, O. Mattelaer, T. Reiter, Comput. Phys. Commun. 183, 1201 (2012). https://doi. org/10.1016/j.cpc.2012.01.022. arXiv:1108.2040 [hep-ph]
- 109. L. Darmé et al., Eur. Phys. J. C 83, 631 (2023). https://doi.org/10. 1140/epjc/s10052-023-11780-9. arXiv:2304.09883 [hep-ph]
- 110. P. Duka, J. Gluza, M. Zralek, Ann. Phys. 280, 336 (2000). https:// doi.org/10.1006/aphy.1999.5988. arXiv:hep-ph/9910279
- 111. A. Roitgrund, G. Eilam, S. Bar-Shalom, Comput. Phys. Commun. 203, 18 (2016). https://doi.org/10.1016/j.cpc.2015.12.009. arXiv:1401.3345 [hep-ph]
- C. Degrande, O. Mattelaer, R. Ruiz, J. Turner, Phys. Rev. D 94, 053002 (2016). https://doi.org/10.1103/PhysRevD.94. 053002. arXiv:1602.06957 [hep-ph]
- 113. B. Fuks, M. Nemevšek, R. Ruiz, Phys. Rev. D 101, 075022 (2020). https://doi.org/10.1103/PhysRevD.101.075022. arXiv:1912.08975 [hep-ph]

- 114. A.A H, B. Fuks, H.-S. Shao, Y. Simon, Phys. Rev. D 107, 075011 (2023). https://doi.org/10.1103/PhysRevD.107.075011. arXiv:2301.03640 [hep-ph]
- 115. O. Mattelaer, M. Mitra, R. Ruiz, (2016). arXiv:1610.08985 [hep-ph]
- 116. B. Fuks, R. Ruiz, JHEP 2017(05), 032 (2017). https://doi.org/10. 1007/JHEP05(2017)032. arXiv:1701.05263 [hep-ph]
- 117. B. Fuks, J. Neundorf, K. Peters, R. Ruiz, M. Saimpert, Phys. Rev. D 103, 115014 (2021). https://doi.org/10.1103/PhysRevD. 103.115014. arXiv:2012.09882 [hep-ph]
- 118. J.C. Romao, J.P. Silva, Int. J. Mod. Phys. A 27, 1230025 (2012). https://doi.org/10.1142/S0217751X12300256. arXiv:1209.6213 [hep-ph]
- S. Dawson, P.P. Giardino, S. Homiller, Phys. Rev. D 103, 075016 (2021). https://doi.org/10.1103/PhysRevD.103.075016. arXiv:2102.02823 [hep-ph]
- 120. M. Nemevšek, F. Nesti, J. Kriewald, LRSM tools page (2024). https://sites.google.com/site/leftrighthep/1-lrsm-feynrules
- M. Nemevšek, F. Nesti, J. Kriewald, LRSM FeynRules model file, version 1.0 NLO (2024). https://feynrules.irmp.ucl.ac.be/wiki/ LRSM_NLO
- W. Grimus, L. Lavoura, JHEP 2000(11), 042 (2000). https://doi. org/10.1088/1126-6708/2000/11/042. arXiv:hep-ph/0008179
- 123. E.K. Akhmedov, W. Rodejohann, JHEP **2008**(06), 106 (2008). https://doi.org/10.1088/1126-6708/2008/06/106. arXiv:0803.2417 [hep-ph]
- 124. M.E. Peskin, D.V. Schroeder, An Introduction to Quantum Field Theory (Addison-Wesley, Reading, 1995)
- 125. T.-P. Cheng, L.-F. Li, *Gauge Theory of Elementary Particle Physics* (Oxford University Press, Oxford, 1984)
- A. Belyaev, N.D. Christensen, A. Pukhov, Comput. Phys. Commun. 184, 1729 (2013). https://doi.org/10.1016/j.cpc.2013.01. 014. arXiv:1207.6082 [hep-ph]

- 127. S. Frixione, B. Fuks, V. Hirschi, K. Mawatari, H.-S. Shao, P.A. Sunder, M. Zaro, JHEP **2019**(12), 008 (2019). https://doi.org/10. 1007/JHEP12(2019)008. arXiv:1907.04898 [hep-ph]
- 128. T. Hahn, Comput. Phys. Commun. **140**, 418 (2001). https://doi. org/10.1016/S0010-4655(01)00290-9. arXiv:hep-ph/0012260
- 129. C. Degrande, Comput. Phys. Commun. **197**, 239 (2015). https:// doi.org/10.1016/j.cpc.2015.08.015. arXiv:1406.3030 [hep-ph]
- J. Alwall, R. Frederix, S. Frixione, V. Hirschi, F. Maltoni, O. Mattelaer, H.S. Shao, T. Stelzer, P. Torrielli, M. Zaro, JHEP 2014(07), 079 (2014). https://doi.org/10.1007/ JHEP07(2014)079. arXiv:1405.0301 [hep-ph]
- 131. A. Buckley, J. Ferrando, S. Lloyd, K. Nordström, B. Page, M. Rüfenacht, M. Schönherr, G. Watt, Eur. Phys. J. C (2015). https://doi.org/10.1140/epjc/s10052-015-3318-8. arXiv:1412.7420 [hep-ph]
- 132. R.D. Ball et al. (NNPDF), Eur. Phys. J. C 82, 428 (2022). https:// doi.org/10.1140/epjc/s10052-022-10328-7. arXiv:2109.02653 [hep-ph]
- 133. A. Hoger, D.E. Carlson, Q. Appl. Math. 42, 113 (1984). https:// doi.org/10.1016/0898-1221(89)90240-X
- 134. L.P. Franca, Comput. Math. Appl. 18, 459 (1989)
- 135. M.A.M. Ashkartizabi, A. Mohammadpour, Calcolo **54**, 643 (2017). https://doi.org/10.1007/s10092-016-0202-3